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# Ophicarbonates evolution from seafloor to subduction and implications for deep-Earth C cycling

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## 1 Abstract

2 The chemical and physical processes operating during subduction-zone metamorphism  
3 can profoundly influence the cycling of elements on Earth. Deep-Earth carbon (C) cycling  
4 and mobility in subduction zones has been of particular recent interest to the scientific  
5 community. Here, we present textural and geochemical data (C-O, Sr isotopes and bulk and  
6 *in-situ* trace element concentrations) for a suite of ophicarbonates (carbonate-bearing  
7 serpentinites) metamorphosed over a range of peak pressure-temperature (*P-T*) conditions  
8 together representing a prograde subduction zone *P-T* path. These rocks, in order of  
9 increasing peak *P-T* conditions, are the Internal Liguride ophicarbonates (from the Bracco  
10 unit, N. Apennines), pumpellyite- and blueschist-facies ophicarbonates from the Voltri Massif  
11 (W. Ligurian Alps) and the Queyras (W. Alps), respectively, and eclogite-facies  
12 ophicarbonates from the Voltri Massif. The Bracco oceanic ophicarbonates retain breccia-like  
13 textures associated with their seafloor hydrothermal and sedimentary origins. Their trace  
14 element concentrations and  $\delta^{18}\text{O}_{\text{VSMOW}}$  (+15.6 to +18.2‰),  $\delta^{13}\text{C}_{\text{VPDB}}$  (+1.1 to +2.5‰) and  
15 their  $^{87}\text{Sr}/^{86}\text{Sr}$  (0.7058 to 0.7068), appear to reflect equilibration during Jurassic seawater-rock  
16 interactions. Intense shear deformation characterizes the more deeply subducted  
17 ophicarbonates, in which prominent calcite recrystallization and carbonation of serpentinite  
18 clasts occurred. The isotopic compositions of the pumpellyite-facies ophicarbonates overlap  
19 those of their oceanic equivalents whereas the most deformed blueschist-facies sample shows  
20 enrichments in radiogenic Sr ( $^{87}\text{Sr}/^{86}\text{Sr} = 0.7075$ ) and depletion in  $^{13}\text{C}$  (with  $\delta^{13}\text{C}$  as low as -  
21 2.0‰). These differing textural and geochemical features for the two suites reflect interaction  
22 with fluids in closed and open systems, respectively. The higher-*P*-metamorphosed  
23 ophicarbonates show strong shear textures, with coexisting antigorite and dolomite, carbonate  
24 veins crosscutting prograde antigorite foliation and, in some cases, relics of magnesite-  
25 nodules enclosed in the foliation. These rocks are characterized by lower  $\delta^{18}\text{O}$  (+10.3 to  
26 13.0‰), enrichment in radiogenic Sr ( $^{87}\text{Sr}/^{86}\text{Sr}$  up to 0.7096) and enrichment in incompatible

27 and fluid-mobile element (FME; e.g., As, Sb, Pb). These data seemingly reflect interaction  
28 with externally-derived metamorphic fluids and the infiltrating fluids likely were derived  
29 from dehydrating serpentinites with hybrid serpentinite-sediment compositions. The  
30 interaction between these two lithologies could have occurred prior to or after dehydration of  
31 the serpentinites. We suggest that decarbonation and dissolution/precipitation processes  
32 operating in ancient subduction zones, and resulting in the mobilization of C, are best traced  
33 by a combination of detailed field and petrographic observations, C, O and Sr isotope  
34 systematics (i.e., 3D isotopes), and FME inventories. Demonstration of such processes is key  
35 to advancing our understanding of the influence of subduction zone metamorphism on the  
36 mobilization of C in subducting reservoirs and the efficiency of delivery of this C to depths  
37 beneath volcanic arcs and into the deeper mantle.

38

39 *Keywords: oceanic ophicarbonates; high-pressure ophicarbonates; deep Carbon cycle;*  
40 *subduction zone; C-O-Sr isotopes*

41

## 42 **1. Introduction**

43 Ophicarbonates (i.e., carbonated ultramafic rocks) are potentially important carbon (C)  
44 reservoirs influencing the long-term C cycle in the solid Earth. Large volumes of  
45 ophicarbonate form by peridotite alteration at the seafloor, causing low- to moderate-  
46 temperature ( $T$ ) serpentinization (e.g., Mével, 2003; O’Hanley, 1996) and carbonation by  
47 mixed seawater and alkaline hydrothermal fluids (e.g., Lost City hydrothermal fields, Früh-  
48 Green et al., 2004; Kelley et al., 2005; Ludwig et al., 2006). Ophicarbonates display  
49 brecciated textures consisting of serpentinite clasts enclosed by sets of carbonate veins and/or  
50 by a micrite matrix. Such rocks occur in present-day oceans (e.g., the Atlantic; Bonatti, 1976)  
51 and their abundance in ophiolitic complexes documents ophicarbonate formation also in  
52 ancient oceans (e.g., Früh-Green et al., 1990; Schwarzenbach et al., 2013; Tartarotti et al.,

53 2019; Weissert and Bernoulli, 1984). Ophicarbonates can also form at depth in subduction  
1 54 zones via interaction of (meta)serpentinite with carbonic fluids leading to carbonation of rock-  
2  
3 55 forming silicates and to C sequestration into oceanic slabs and supra-subduction mantle  
4  
5  
6 56 (Piccoli et al., 2016; Scambelluri et al., 2016). These two processes together affect C storage  
7  
8 57 and transport from oceans to subduction zones and the transfer of C-O-H fluids into volcanic  
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10  
11 58 arcs (Alt et al., 2013; Collins et al., 2015; Kerrick and Connolly, 1998; Scambelluri et al.,  
12  
13 59 2016) or in deeper parts of the mantle.

14  
15 60 Experiments and thermodynamic modelling suggest relative immobility of C in  
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17  
18 61 subducting oceanic crustal rocks to very high pressure-temperatures ( $P$ - $T$ ) conditions unless  
19  
20 62 the rocks are infiltrated by H<sub>2</sub>O-rich fluids (Collins et al., 2015; Kerrick and Connolly, 2001,  
21  
22 63 1998; Molina and Poli, 2000; Poli et al., 2009). Studies of exposed high- $P$ / $T$  metamorphic  
23  
24  
25 64 rocks document some C mobilization during subduction either by decarbonation reactions or  
26  
27 65 by carbonate dissolution in aqueous fluids (Ague and Nicolescu, 2014; Cook-Kollars et al.,  
28  
29 66 2014; Frezzotti et al., 2011; Malaspina et al., 2009; Piccoli et al., 2016; Sapienza et al., 2009;  
30  
31  
32 67 Scambelluri et al., 2016); however, the scale of this related C transport has remained unclear  
33  
34  
35 68 (see the discussion by Epstein et al. 2019). Field and stable isotope studies of high- $P$  and  
36  
37 69 ultrahigh-pressure ( $UHP$ ) metamorphic suites indicate that the extent of C mobilization by  
38  
39  
40 70 these processes is highly dependent upon the degree to which rocks behave as systems open  
41  
42 71 to infiltration by externally-derived H<sub>2</sub>O-rich fluids (Collins et al., 2015; Epstein et al., 2019).

43  
44 72 Serpentinite is pivotal in the subduction zone C cycle because, as the result of antigorite  
45  
46  
47 73 dehydration, it supplies H<sub>2</sub>O to adjacent carbonate rocks, thus potentially driving  
48  
49 74 decarbonation reactions (i.e., destabilization of carbonates) and carbonate dissolution.  
50  
51  
52 75 Moreover, serpentine is reactive to C-O-H fluids and sequesters CO<sub>2</sub> from fluids via  
53  
54 76 carbonation of its silicate minerals (Scambelluri et al., 2016). The study of ophicarbonates is  
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57 77 timely, given recent interest in subduction zone C cycling, and can address the mechanisms of  
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78 C-storage and loss within a single, complex rock system and between serpentinite and  
79 carbonate-bearing reservoirs.

80 Abundant petrographic and petrologic work available to date deals with present-day and  
81 fossil oceanic ophicarbonates (Bonatti, 1976; Clerc et al., 2014; Collins et al., 2015;  
82 Cortesogno et al., 1980; Driesner, 1993; Galli and Togliatti, 1965; Lafay et al., 2017; Treves  
83 et al., 1995; Weissert and Bernoulli, 1984); however, less attention has been paid to  
84 ophicarbonate evolution during subduction-zone metamorphism (Collins et al., 2015; Debret  
85 et al., 2018; Driesner, 1993; Scambelluri et al., 2016). We focused on ophicarbonates exposed  
86 in Alpine-Apennine ophiolite sequences representing a wide range of peak metamorphic  $P$ - $T$ ,  
87 from low- $T$  oceanic conditions recorded by Apennine ophicarbonates, to pumpellyite-,  
88 blueschist- and eclogite-facies conditions experienced by ophicarbonates from the Ligurian  
89 and the Western Alps. In this paper, based on study of high- $P$  metamorphic rocks, we  
90 demonstrate that decarbonation and dissolution/precipitation processes operating in ancient  
91 subduction zones are best traced by a combination of detailed field and petrographic  
92 observations, FME inventories, and C, O, and Sr isotope systematics (i.e., 3D isotopes).

93

## 94 **2. Geological setting**

95

### 96 *2.1. Northern Ligurian Apennine ophicarbonates*

97 In the Northern Ligurian Apennine, ophiolites either occur as large olistoliths (in the  
98 sedimentary flysch of the External Liguride Units) or form large coherent bodies (in the  
99 Internal Liguride Units). In the latter setting, ophicarbonates (**Fig. 1A**) occur at the  
100 stratigraphic top of the serpentinitized mantle (e.g., Cortesogno et al., 1980) and are overlain by  
101 pillow basalt and oceanic sediments (i.e., radiolarian cherts, limestone and mudstone). In such  
102 a cases, the Moho is commonly represented by preserved primary contacts of serpentinite-  
103 ophicarbonate with pillow basalt and, locally, with extensional tectonic breccia and deep

104 oceanic sediments (Alt et al., 2018; Decandia et al., 1998; Lagabrielle, 1987). Thus far,  
1 105 tectonic (OCI) and sedimentary (OCII) ophiicarbonates have been distinguished based on  
2  
3 106 textural and petrographic features (e.g., Lemoine et al., 1987). The OCI type is characterized  
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5  
6 107 by polyphase brittle-ductile to brittle deformation (Treves and Harper, 1994). These rocks  
7  
8 108 display early serpentine formation after mantle minerals and in shear extensional veins,  
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10  
11 109 followed by infiltration of CO<sub>2</sub>-rich fluids to form a sequence of generations of calcite veins  
12  
13 110 in response to increasing hydrothermal fluid pressure ( $P_{\text{fluid}}/P_{\text{lithostatic}} > 1$ ). The last brittle  
14  
15 111 deformation led to formation of veins filled by calcite and talc druses (Treves and Harper,  
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17  
18 112 1994). The set of tectonic and hydrothermal structures recorded by the OCI is related to the  
19  
20 113 tectonic exposure of the oceanic mantle at surface levels under progressively lower  
21  
22  
23 114 temperature conditions at extensional core-complexes at slow-spreading ridges. The OCI  
24  
25 115 samples are thus representative of extensional fault settings (Treves and Harper, 1994) and are  
26  
27  
28 116 comparable with ophiicarbonates beneath the Lost City hydrothermal vent at the Mid-Atlantic  
29  
30 117 Ridge (e.g., Ludwig et al., 2006). Tectonic-sedimentary reworking of the exposed  
31  
32  
33 118 serpentinized mantle, coupled with calcite precipitation, leads to the formation of OCII  
34  
35 119 showing angular to sub-angular serpentinite clasts of varying sizes embedded in a micrite  
36  
37 120 sedimentary matrix (Treves and Harper, 1994). Oxidation and carbonation processes  
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40 121 accompany the hydrothermal alteration, allowing formation of hematite and thus conferring  
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42 122 upon these rocks a typical red color during a process thought to have generated the Ligurian  
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44  
45 123 ophi carbonate variety known as *Rosso di Levante* (Galli, 1957).

46  
47 124

## 48 49 125 *2.2 Western Alps*

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51  
52 126 The Alpine architecture consists of large tectonic units of continental and oceanic  
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54 127 affinity (e.g., Dal Piaz et al., 2003): the ophi carbonate rocks studied here belong to the  
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56  
57 128 oceanic lithosphere exposed at three main localities of the Upper Penninic nappe stack and  
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59 129 recording progressively higher peak metamorphic  $P$ - $T$  during subduction (**Fig. 1**). From lower  
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130 to higher grade, the sampled localities are: the pumpellyite-actinolite facies Sestri-Voltaggio  
1 131 Zone (Ligurian Alps), the blueschist-facies Queyras (Ubaye locality; French Alps), and the  
2  
3 132 eclogite-facies-metamorphosed Voltri Massif (Ligurian Alps).  
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6 133

#### 8 134 *2.2.1 Sestri-Voltaggio zone*

10 135 The Sestri-Voltaggio zone consists of several tectonic units and is believed to represent  
11  
12  
13 136 the N-S junction zone between the Alpine chain and the Northern Apennines (Crispini and  
14  
15 137 Capponi, 2001). The ophicarbonates rocks sampled in this area belong to the Figogna unit,  
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17  
18 138 consisting of oceanic lithospheric mantle comparable to that in the Voltri Massif, but which  
19  
20 139 escaped eclogite facies metamorphism during Alpine subduction. The metamorphic grade of  
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22  
23 140 this unit is pumpellyite-actinolite facies, in the stability field of lawsonite (270-320 °C and 0.6  
24  
25 141 GPa; Leoni et al., 1996). Here ophicarbonates crop out near the village of Pietralavezzara in  
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27  
28 142 decametric-sized slices with N-S orientations (PL samples in **Fig. 1C**). Despite the polyphase  
29  
30 143 deformation that affected the Figogna unit, the original oceanic sequence showing  
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32  
33 144 ophicarbonates on top of the ultramafic and gabbroic basement, and below a volcano-  
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35 145 sedimentary cover, can still be recognized (Crispini and Capponi, 2001).  
36

37 146

#### 40 147 *2.2.2 Queyras complex*

42 148 The Queyras complex is part of the Piedmont-Ligurian domain, west of the eclogite-  
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44  
45 149 facies Monviso ophiolite and near the contact with the Briançonnais domain (**Fig. 1B**). This  
46  
47 150 complex was interpreted as a sedimentary accretionary wedge composed of oceanic sediments  
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49  
50 151 and accreted slices of oceanic lithosphere (Lafay et al., 2013; Lagabrielle et al., 1985). The  
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52 152 metamorphic grade increases eastwards, from temperature < 350 °C and pressure < 1.1 GPa,  
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54  
55 153 to *P-T* conditions pertaining to the blueschist-facies conditions (up to 480 °C and  $P > 1.2$   
56  
57 154 GPa; Agard et al., 2001). The Queyras ophiolite crops out as scattered large olistoliths within  
58  
59 155 the metasedimentary Schistes Lustrés: the ophiolite bodies preserve the original internal  
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156 structure and stratigraphy, enabling reconstruction of their oceanic history (Tricart and  
1 157 Lemoine, 1991). Within each body, three major components of the oceanic lithosphere can be  
2  
3 158 distinguished; from bottom to top these are (i) an ultramafic sequence of serpentinitized  
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6 159 mantle-derived peridotites, locally crosscut by Late Triassic gabbro intrusions (Carpena and  
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8 160 Caby, 1984), (ii) a discontinuous basaltic layer of tholeiitic pillow basalts, and (iii) a  
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11 161 sedimentary pelagic cover represented by the Schists Lustrés complex. Where the basaltic  
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13 162 sequence is missing, the ultramafic and the sedimentary rocks can be directly in contact  
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16 163 (Tricart and Lemoine, 1991). The ophicarbonates are on top of the ultramafic sequence  
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18 164 and can be subdivided into groups of tectonic (OCI) and sedimentary (OCII) origin. The first  
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20  
21 165 type is characterized by a dense network of calcite veins and fractures formed by tectonic  
22  
23 166 extension and hydraulic fracturing; the OCII type derives from in-situ reworking of OCI, of  
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25 167 serpentinites and/or basaltic materials forming clasts in a micrite matrix (e.g., Lemoine et al.,  
26  
27  
28 168 1987). The presence of ophicarbonates is interpreted as representing exposure of mantle  
29  
30 169 to seawater on the ocean floor (Tricart and Lemoine, 1991). The ophicate samples  
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33 170 analyzed in this study (CU samples in **Fig. 1B**) experienced blueschist-facies *P-T* conditions  
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35 171 (325 – 375 °C and 1.1 – 1.4 GPa; Michard et al., 2004).

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### 40 173 *2.2.3 Voltri Massif*

42 174 The ophiolitic Voltri Massif is a blueschist- to eclogite-facies complex of Jurassic  
43  
44  
45 175 oceanic lithosphere composed of serpentinites hosting gabbro- and basalt-derived eclogite and  
46  
47 176 metarodingite. This sequence is associated with high-*P*-metamorphosed oceanic sediment, i.e.  
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50 177 calc- and mica-schists (referred to as Schistes Lustrés) interlayered with metabasalts and  
51  
52 178 cherts (e.g., Chiesa et al., 1975). Remnants of sub-continental lithospheric mantle are well-  
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54  
55 179 exposed in the Erro-Tobbio metaperidotite unit, representing the pre-Alpine rifting and ocean-  
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57 180 floor alteration followed by subduction and metamorphism at eclogite-facies *P-T* conditions  
58  
59 181 (Hermann et al., 2000; Hoogerduijn Strating et al., 1993; Scambelluri et al., 1995, 1991;  
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182 Visser et al., 1995, 1991). Peak metamorphic  $P$ - $T$  estimates for the Voltri Massif are about  
1 183 450 – 500 °C and 1.3 – 1.7 GPa (Cimmino et al., 1979; Federico et al., 2004; Messiga and  
2  
3 184 Scambelluri, 1991), conditions that produced metamorphic olivine + antigorite in  
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6 185 serpentinites, garnet + omphacite + rutile in eclogite and high-Si-bearing phengite + garnet in  
7  
8 186 metasedimentary rocks. Field and geochemical evidence (Cannaò et al., 2016; Federico et al.,  
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10  
11 187 2007; Scambelluri and Tonarini, 2012), together with numerical modelling (Malatesta et al.,  
12  
13 188 2012), suggests a complex history for the Voltri Massif during prograde and retrograde  
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16 189 evolution along the slab-mantle interface.

17  
18 190 The ophicarbonates of the Voltri Massif crop out in several localities (**Fig. 1**) and  
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20  
21 191 were first described by Cortesogno et al. (1981, 1980). These rocks consist of carbonate  
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23 192 mineral phases (calcite, dolomite and/or magnesite) coexisting with antigorite and, locally,  
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25 193 metamorphic olivine. These rocks are interpreted as the metamorphic counterparts in the non-  
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27  
28 194 subducted oceanic ophicarbonates of the Northern Apennines ophiolites, with grades ranging  
29  
30 195 from greenschist-, to blueschist-, to eclogite-facies. Recently, Scambelluri et al. (2016)  
31  
32  
33 196 described in detail the evolution, metamorphic reactions, and C and O isotope compositions at  
34  
35 197 an outcrop within the Voltri Massif, highlighting the high potential of serpentinite for  
36  
37  
38 198 sequestering C and thus influencing the deep C cycle. The samples from the Voltri Massif  
39  
40 199 analyzed in this study are from two localities representing ophicarbonates affected by the  
41  
42 200 eclogite-facies metamorphism: La Pesca (Scambelluri et al., 2016; LP13 EOC samples) and  
43  
44  
45 201 Sant'Anna (SA samples).

### 50 203 **3. Petrography and microstructures**

51  
52 204 The ophicate samples presented here retain evidence of complex polyphase  
53  
54 205 histories. Here, we describe the key features of representative ophicate samples from  
55  
56  
57 206 each locality: for the Northern Apennine, we retain the classification of the ophicarbonates of  
58  
59 207 tectonic-hydrothermal (OCI) and sedimentary (OCII) origin initially proposed by Lemoine et  
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208 al. (1987). Also, we describe the oceanic serpentinite exposed in the Northern Apennine,  
1 209 regarding it as the starting material (protoliths) for the higher-grade ophicarbonates.  
2

3 210  
4

### 6 211 ***3.1 Oceanic Northern Apennine Ophicarbonates***

#### 8 212 *Serpentinite protolith rocks*

10 213 The ultramafic protoliths of the oceanic ophicarbonates are represented by highly (90%)  
11  
12  
13 214 to fully (100%) serpentinitized mantle peridotites. In these rocks, static pseudomorphic  
14  
15 215 replacement led to formation of mesh serpentine + magnetite textures after mantle olivine and  
16  
17  
18 216 bastite after mantle pyroxenes in which traces of the former pyroxene cleavage are still  
19  
20 217 recognizable. The medium grain size of the pseudomorphs and the isotropic texture suggest  
21  
22  
23 218 an original harzburgite for these ultramafic rocks. The sample collected in this study does not  
24  
25 219 display evidence for deformation, but high-*T* shear domains are described in the literature for  
26  
27  
28 220 the serpentinite basement (Treves and Harper, 1994).  
29

30 221  
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#### 32 222 *Ophicarbonates of tectonic and hydrothermal origin (OCI)*

34  
35 223 The ophicarbonates of tectonic and hydrothermal origin (OCI) contain several  
36  
37 224 generations of carbonate veins overprinting the serpentinitized precursor. Calcium carbonate is  
38  
39  
40 225 the most widespread mineral, ubiquitous in these rocks at this stage. Samples can display  
41  
42 226 variably extensive carbonation of the starting serpentinite during the hydrothermal activity.  
43  
44  
45 227 The amounts of vein-related calcite can range from few modal % calcite (e.g., ZRL 15-3) to  
46  
47 228 more than 80 % (e.g., MNT 13-5) of the host rock. The host serpentinite has the texture of a  
48  
49  
50 229 jigsaw puzzle breccia with clasts ranging in size from the outcrop scale (m in size) to hand-  
51  
52 230 sample size (few cm)- and to mm-scale. The ultramafic clasts also display varying extents of  
53  
54 231 carbonation, from less than 20% (ISC 13-1) to 90% (ZRL 13-1) in volume.  
55  
56

57 232 The ultramafic clasts correspond to fully serpentinitized peridotites, which do not  
58  
59 233 preserve olivine and which rarely preserve relict mantle pyroxene and spinel partially  
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234 converted to magnetite. The clasts of the host serpentinite are heterogeneous in size and  
1 235 polygonal in shape, with sub-angular to angular edges. Each clasts preserves a record of the  
2  
3 236 original undeformed and deformed structure of the serpentinitized mantle peridotite host. The  
4  
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6 237 monotonous chrysotile + magnetite assemblage of the pure serpentinite clasts is locally  
7  
8 238 accompanied by formation of chlorite, talc and tremolite. Near the calcite hydrothermal veins,  
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11 239 the serpentinite clasts unaffected by carbonation and preserving mantle pyroxene relics show  
12  
13 240 evidence of infiltration in the form of small- to medium- sized calcite veinlets.

14  
15 241 The serpentinite clasts are affected by varying degrees of carbonation: from minor  
16  
17  
18 242 overgrowth by neoblastic fine-grained calcite to reddish clasts showing pervasive to full  
19  
20 243 carbonation associated with oxidation. This carbonation represent a key feature of such  
21  
22  
23 244 hydrothermally-altered oceanic ophicarbonates: full carbonation of the ultramafic silicate  
24  
25 245 clasts concurs with formation of variably abundant hematite after magnetite. Replacement by  
26  
27  
28 246 calcite of serpentine from the core of the mesh texture and along cleavage of the former  
29  
30 247 pyroxene by calcite occurs in nearly all oceanic samples. Locally, along the cleavages of  
31  
32  
33 248 serpentinitized pyroxenes, trails of small green andraditic garnet crystals are intergrown with  
34  
35 249 serpentine and calcite (**Fig. 2C**). In samples where the carbonation has been quite extensive,  
36  
37  
38 250 the serpentinite clasts are entirely replaced by fine-scale calcite grains and a film of hematite  
39  
40 251 grows along the border, giving them a ghost-like texture (**Fig. 2A, B**). Interestingly, the full  
41  
42 252 carbonation of the ultramafic silicate clasts appears to be associated with the formation of  
43  
44  
45 253 abundant hematite (**Fig. 2A, B**), suggesting that the  $\text{Fe}^{2+}$  released by the carbonated silicates  
46  
47 254 was converted to  $\text{Fe}^{3+}$  thus stabilizing hematite due to increased oxidizing conditions.

48  
49 255 The last brittle deformation event recorded by these rocks is represented by the  
50  
51  
52 256 formation of calcite veins that crosscut all previous structures. Calcite is coarse-grained and  
53  
54  
55 257 shows euhedral habit and, typically, twinning. Talc is observed in drusy textures commonly in  
56  
57 258 the core of the veins, seemingly representing as the last precipitates. In some samples, calcite  
58  
59 259 is replaced by talc aggregates (**Fig. 2D**).

1 261 *Sedimentary ophicarbonates (OCII)*

2  
3 262 These rocks display a sedimentary breccia-like texture with clasts consisting of  
4  
5  
6 263 serpentinites or of reworked OCI. The sedimentary OCII clasts typically display angular  
7  
8 264 edges and widely varying grain size from several millimeters to few hundreds of micrometers.  
9  
10  
11 265 The mineral assemblage of the clasts, as in the OCI, consists of serpentine, magnetite, chlorite,  
12  
13 266 talc and minor tremolite. Based on the extent of carbonation and oxidation, two types of  
14  
15 267 sedimentary breccias can be identified: (i) samples SG 15-4 and MNT 15-2, showing minor  
16  
17  
18 268 and abundant carbonation and oxidation, respectively, and (ii) sample SG 15-2, showing  
19  
20 269 intermediate amounts of carbonation and oxidation.

21  
22  
23 270 The green colour of ophicarbonate SG 15-4 reflects the high modal abundance of  
24  
25 271 hydrous silicate minerals still preserved in the rock. The textures of the serpentinite clasts  
26  
27  
28 272 preserve the mantle mineral domains; however, some clasts with complex texture likely  
29  
30 273 derive from pre-existing serpentine veins. The matrix calcite range from micritic to sparry, the  
31  
32  
33 274 latter occurrence type showing twinning. The sedimentary structures are cut by calcite veins  
34  
35 275 showing, from rim to core, a change from botryoidal to irregular blocky textures (**Fig. 2E**).

36  
37 276 The ophicarbonate sample MNT 15-2 shows an intense degree of carbonation and  
38  
39  
40 277 oxidation, with a reddish color due a large amount of hematite (**Fig. 2F**). In sample SG 15-2,  
41  
42 278 only a few % of the clasts still preserve the original serpentinite composition, while  
43  
44  
45 279 widespread ghost-like textures derive from full replacement of former silicate minerals by  
46  
47 280 fine-grained calcite. Locally, the clasts are composed of single crystals of calcite, probably  
48  
49  
50 281 derived from reworking of the OCI and associated hydrothermal calcite veins. The matrix of  
51  
52 282 sample MNT 15-2 consists mainly of medium- to fine-scale (to microcrystalline) grains of  
53  
54 283 anhedral calcite with lobate and irregular intercrystalline boundaries. Along the contact  
55  
56  
57 284 between the clasts and the matrix, calcite grows as cement crusts typically forming  
58  
59 285 gravitational beard-like patterns.  
60  
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63  
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65

### 3.2 Subduction-related ophicarbonates

#### *Pumpellyite-actinolite-facies ophicarbonates*

Samples PL 13-1 and PL 13-2 display two texturally different types of micro-domains indicating multiple stages of recrystallization of calcite grains. The first micro-domain consists of coarse- to medium-sized grains of calcite (1) showing development of a second generation of twin lamellae cutting at high angles the earlier-formed generation and affected by growth of twin-free calcite grains. The coarse calcite (1) is cut by shear zones that represent the second type of micro-domain observed in these samples. The shear zones consist of medium- to fine- grained, un-twinned calcite (2) irregular in shape and with complex grain boundaries. Along the shear zones, calcite (1) is partially recrystallized and becomes elongated and re-oriented coherently into the shear zone orientation (**Fig. 3A**). Both samples (PL 13-1 and PL 13-2) also contain 10-20 volume % serpentinite clasts ranging in size from several centimetres to a few millimetres. In general, the largest clasts are broken into smaller pieces when embedded and displaced along the shear-zones as sigmoidal porphyroclasts. The serpentinite clasts consist of magnetite and antigorite showing tabular and interlocking textures grown at the expense of mesh-like lizardite/chrysotile. In shear zones, thin dark films of antigorite are interlayered with the calcite (2), giving to the rock an S-C like texture. The serpentinite clasts in these samples thus derive from an early stage of oceanic serpentinitization and carbonation.

A key feature pertaining to these samples is the replacement of antigorite within the clasts by neoblastic calcite. Calcite overgrowths occur along the contacts between antigorite and calcite within the shear zones and inside the various serpentinite clasts (**Fig. 3B**). These textures indicate that replacement of serpentine by calcite was due to syn-deformation infiltration by carbonic fluids.

312 *Blueschist-facies ophicarbonates*

1 313 Samples CU3 and CU 13-1 represent ophicarbonates re-equilibrated at blueschist-facies  
2  
3 314 *P-T* conditions. CU3 consists of coarse serpentinite clasts in a coarse-grained calcite matrix.  
4  
5  
6 315 In the clasts, antigorite overgrows earlier serpentine minerals and, in the majority of the clasts,  
7  
8 316 the original serpentinite texture is cross-cut and overgrown by networks of calcite-bearing  
9  
10  
11 317 veins. The coarse-grained matrix-forming calcite is anhedral and shows corroded grain  
12  
13 318 boundaries, undulose extinction, and abrupt change in the twin growth, features that together  
14  
15 319 represent deformation during crystallization. The calcite grains formed along the contacts  
16  
17  
18 320 with the serpentinite clasts and within the fracture network (**Fig. 3C**) display elongate textures  
19  
20 321 suggesting syn-deformation crystallization.

22  
23 322 Sample CU 13-1 shows the development of carbonate-rich shear zones alternating with  
24  
25 323 domains with abundant serpentinite clasts. The shear zone textures are quite similar to those  
26  
27  
28 324 in samples PL 13-1 and PL 13-2 and are characterized by re-orientation of calcite and  
29  
30 325 serpentinite clasts along the shear foliation and by crystallization of neoblastic calcite  
31  
32 326 emanating from the grain boundaries of former calcite (**Fig. 3D**). Deformation within the  
33  
34  
35 327 shear zones leads to fracturing of the serpentinite clasts, grain size reduction, and uniform  
36  
37 328 orientation of the antigorite flakes. Within the shear zones, the matrix calcite is very fine-  
38  
39  
40 329 grained (maximum 100  $\mu\text{m}$ ) and shows heterogranular and hypidiotopic fabric typical of  
41  
42 330 dynamic recrystallization and imparting a marble-like texture. The lack of twin lamellae in  
43  
44  
45 331 this type of calcite suggests an increase in the *T* conditions with respect to that experienced by  
46  
47 332 the coarse-grained calcite crystals.

48  
49  
50 333

51  
52 334 *Eclogite-facies ophicarbonates*

53  
54 335 Samples from the La Pesca locality (LP13 EOC1 and LP13 EOC3) display an antigorite  
55  
56  
57 336 + calcite + magnetite ( $\pm$  hydro-garnet in sample LP13 EOC1) mineral assemblage. As in the  
58  
59 337 oceanic samples, garnet forms along the cleavages of the serpentinitized pyroxene and, locally,  
60  
61  
62  
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64  
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338 around magnetite pods displaced along the main foliation (**Fig. 3E**). As often occurs in  
1 339 serpentinite, the magnetite pods may derive from oceanic alteration and recrystallization of  
2  
3 340 Cr-rich mantle spinel. The main difference between the two La Pesca samples is the strong  
4  
5 341 deformation affecting LP13 EOC3, a texture similar to that of samples CU 13-1 and PL 13-  
6  
7  
8 342 1/2. Despite the textural similarity to CU 13-1 and PL 13-1/2, deformed calcite in sample  
9  
10 343 LP13 EOC3 is much finer-grained and contains more equigranular fabrics and hypidiotopic to  
11  
12 344 idiotopic crystals. Sample LP13 EOC1 largely consists of calcite veins cutting the entire rock  
13  
14  
15 345 structure and suggesting late (subduction-related?) infiltration of C-bearing fluids along  
16  
17 346 fractures.

20 347 The S. Anna ophicarbonates are the only samples showing dolomite and magnesite  
21  
22 348 associated with antigorite, magnetite, minor chlorite, and diopside. In these rocks, coarse  
23  
24  
25 349 porphyroclastic magnesite occurs within the foliation produced at high-*P* conditions, probably  
26  
27 350 as the result of tectonic dismemberment of former veins (sample SA 15-8). An interesting  
28  
29  
30 351 feature of this magnesite is the presence of primary and pseudo-secondary fluid inclusions  
31  
32 352 (**Figs. 3G, 3H**) showing tabular shapes, variable size (from 20 to few  $\mu\text{m}$ ) and infilling by  
33  
34  
35 353 liquid + gas. Occasionally, small aggregates of Ti-clinohumite are present and associated with  
36  
37 354 the magnesite grains. Magnesite is partially replaced by dolomite and calcite along contacts  
38  
39  
40 355 with antigorite and, concurrent with this replacement, acicular tremolite was formed after  
41  
42 356 antigorite.

45 357 Sample SA 15-1 shows micro-scale shear zones similar to those found in the other  
46  
47 358 subduction-related samples: in such sheared domains, dolomite is overgrown by calcite,  
48  
49 359 antigorite and minor chlorite likely related to retrograde deformation. Though high-*P*  
50  
51  
52 360 conditions were reached in the S. Anna area, as described in literature, no evidence of  
53  
54 361 metamorphic olivine has been found in the samples investigated here. However, we interpret  
55  
56  
57 362 the dolomite + antigorite + diopside assemblage (**Fig. 3F**) as representing high-*P* conditions

363 comparable with those of the ophicarbonates from the Zermatt-Saas Unit (Chatillon area, W.  
1 364 Alps), the latter peak-metamorphosed at 1.4 GPa and 450-550 °C (Driesner, 1993).

2  
3 365

#### 4 366 **4. Methods**

##### 7 8 367 *4.1 Bulk rocks*

9  
10 368 Whole rock major element analyses were performed at ActLabs Ltd. using a fusion  
11  
12  
13 369 technique employing a Li-metaborate/tetraborate fusion digested in a weak nitric acid and  
14  
15 370 measured by ICP-OES. Accuracy and precision are estimated as better than 1%. Loss on  
16  
17  
18 371 ignition (LOI) was determined by weight difference following ignition of samples in a  
19  
20 372 furnace. The amount of CO<sub>2</sub> was determined by calcimetry at the IGG-CNR (Pisa, Italy) by  
21  
22  
23 373 reacting small aliquot of sample (about 0.3 g) with HCl, and measuring the release of CO<sub>2</sub>  
24  
25 374 from carbonates through the volume of water shifted in a graduated cylinder. The  
26  
27  
28 375 concentration is then calculated using a calibration line. The CO<sub>2</sub> measured from sample  
29  
30 376 containing the less-reactive magnesite in the assemblage (i.e., SA 15-8) may be  
31  
32  
33 377 underestimated. The amount of H<sub>2</sub>O was estimated by the simple difference between LOI and  
34  
35 378 CO<sub>2</sub> contents for each sample.

36  
37 379 Trace element concentrations (except for B) were accomplished using an Agilent 7700x  
38  
39  
40 380 quadrupole ICP-MS at Géosciences Montpellier (AETE-ISO facility, Montpellier, France)  
41  
42 381 following the procedure described in Godard et al. (2000) and Ionov et al. (1992). Boron  
43  
44  
45 382 concentrations for all samples were determined by isotope dilution using the NIST SRM 982  
46  
47 383 spike in a single collector VG Isomass 54E positive thermal ionization mass spectrometer (P-  
48  
49  
50 384 TIMS) running in dynamic mode at the IGG-CNR (Pisa, Italy).

51  
52 385

##### 53 54 386 *4.2 Mineral compositions*

55  
56  
57 387 Major element mineral compositions were measured at the Department of Earth  
58  
59 388 Sciences, University of Milano (Italy), using the electron microprobe JEOL Superprobe 8200.

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61  
62  
63  
64  
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389 The analyses were conducted with a WDS system at 15.0 kV, 5 nA for the electron beam and  
1 390 1  $\mu\text{m}$  beam size. We used natural silicate standards (the numbers refer to the international  
2  
3 391 standards): Mg on olivine 153, Fe on fayalite 143, Na on omphacite 154, Ti on ilmenite, Mn  
4  
5  
6 392 on rodonite, K on k-feldspar 113, Al and Ca on anorthite 137, Cr on metallic/pure Cr and Si  
7  
8 393 on wollastonite.

10  
11 394 Mineral trace element concentrations were determined at the University of Bern by laser  
12  
13 395 ablation, using a Geolas Pro 2006 193 nm ArF excimer laser coupled to an Elan DRC-e ICP-  
14  
15 396 MS instrument (LA-ICP-MS). Analytical strategies, data reduction and instrument  
16  
17  
18 397 optimization procedures closely follow those reported by Spandler et al. (2011) and Pettke et  
19  
20 398 al. (2012). Beam sizes were as large as possible (up to 160  $\mu\text{m}$  diameter) to lower the  
21  
22  
23 399 detection limits, calculated using the stringent formulation in Pettke et al. (2012). External  
24  
25 400 standardization was against the GSD-1D basalt glass, employing the values cited in Peters  
26  
27  
28 401 and Pettke (2017) and the GeoRem data base (Jochum et al., 2005). Data reduction employed  
29  
30 402 SILLS (Guillong et al., 2008), and the sum of all measured element oxides (total oxides, i.e.,  
31  
32  
33 403 100 wt.% minus LOI; e.g., Halter et al., 2002) was used as the internal standard for  
34  
35 404 quantification. Due to the lack of data for  $\text{Fe}^{3+}/\text{Fe}^{2+}$ , total Fe was calculated as FeO. Ablation  
36  
37  
38 405 rates were turned to c. 0.08  $\mu\text{m}$  per pulse via attenuation of the laser output beam; the laser  
39  
40 406 repetition rate was 10 Hz. Measurement accuracy and precision were monitored by analysing  
41  
42 407 SRM612 glass from NIST as an unknown, and the data mostly agree to within 5% of the  
43  
44  
45 408 reference values (**Supplementary Table A3**).

#### 49 50 410 *4.3 Isotopic analyses*

51  
52 411 The samples were analyzed for  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  at the Department of Earth and  
53  
54  
55 412 Environmental Sciences, Lehigh University (USA), using a Finnigan MAT 252 gas mass-  
56  
57 413 spectrometer coupled with a GasBench II system, applying the methods described by Cook-  
58  
59 414 Kollars et al. (2014) and Collins et al. (2015). We analyzed bulk rocks as well as more pure  
60  
61  
62  
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415 carbonates from different textural settings sampled by micro-drilling, using tungsten carbide  
1 416 bits with diameter of 1 and 2 mm. Calcite, dolomite or magnesite was reacted at 72 °C for 30  
2  
3 417 min, 3 or 6 hours, respectively, with 0.2 mL 100% phosphoric acid (H<sub>3</sub>PO<sub>4</sub>) in order to  
4  
5 418 release CO<sub>2</sub>, according to the methods of McCrea (1950), Paul and Skrzypek (2007), and  
6  
7  
8 419 Sharma et al. (2002). All data are presented in standard delta(δ)-notation, reported relative to  
9  
10 420 VPDB and VSMOW for δ<sup>13</sup>C and δ<sup>18</sup>O, respectively. Regular analyses of an in-house  
11  
12 421 standard and the international standard NBS-19 allowed monitoring and correction of the data,  
13  
14  
15 422 resulting in a standard deviation (1σ) of ~ 0.2‰ for both δ<sup>13</sup>C and δ<sup>18</sup>O.  
16

17  
18 423 Strontium isotope analyses were performed using a Finnigan MAT 262 multicollector  
19  
20 424 mass-spectrometer (at the IGG-CNR, Pisa, Italy), operated in dynamic mode, after ion-  
21  
22 425 exchange purification through Sr-spec resin. Measured <sup>87</sup>Sr/<sup>86</sup>Sr ratios were normalized to  
23  
24  
25 426 <sup>86</sup>Sr/<sup>88</sup>Sr = 0.1194. Replicate analyses of NIST SRM 987 (SrCO<sub>3</sub>) standard gave an average  
26  
27 427 value of 0.710207 ± 13 (2SD, n = 5). Published values are adjusted to <sup>87</sup>Sr/<sup>86</sup>Sr = 0.710250.  
28  
29 428 Strontium blanks were about 0.3 ng, which are negligible given the Sr concentrations in the  
30  
31  
32 429 samples analyzed in this study.  
33

34  
35 430

## 36 37 431 **5. Results**

### 38 39 432 *5.1 Bulk rock major and trace element concentrations*

40 433 The major element whole-rock concentrations of the samples are reported in **Table 1**.  
41  
42 434 The compositions of ophicarbonates are linked to the proportion of ultramafic and carbonate  
43  
44 435 component present: here, the composition of a pure serpentinite and an oceanic limestone  
45  
46 436 (MNT 15-1 and ZRL 15-7, respectively) were considered as the end-members for all of the  
47  
48 437 ophicarbonate rocks that were investigated. The major element compositions of OCI and  
49  
50 438 OCII ophicarbonates plot along a linear array in a CaO, MgO and SiO<sub>2</sub> diagram, as do the  
51  
52 439 subduction-related ophicarbonates. Similar variations in CO<sub>2</sub> and H<sub>2</sub>O with respect to MgO  
53  
54 440 and CaO are displayed for all ophicarbonates, with rocks showing more evidence of  
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441 recrystallization during subduction processes (e.g. PL 13-2, CU 13-1 and LP13 EOC3)  
1 442 plotting near the CO<sub>2</sub>-rich end-member (**Fig. 4**). Sample SA 15-8, which contains magnesite,  
2  
3 443 has a slightly higher MgO concentration and a lower CaO concentrations than the other  
4  
5  
6 444 subduction-related ophicarbonates (**Fig. 4**).

7  
8 445 Trace element whole-rock compositions of the samples are reported in **Table 1** and in  
9  
10  
11 446 **Fig. 5**. The pure serpentinite sample MNT 15-1 shows low REE concentrations with respect  
12  
13 447 to primitive mantle values (PM, McDonough and Sun, 1995) and a negligible negative Eu  
14  
15 448 anomaly. The strong light (L)REE depletion relative to concentrations of medium (M)REE  
16  
17  
18 449 and heavy (H)REE ( $La_N/Nd_N = 0.15$ ;  $Sm_N/Ho_N = 0.36$ ;  $Er_N/Lu_N = 0.81$ ) suggests a depleted  
19  
20 450 harzburgite protolith. The limestone sample ZRL 15-7 is enriched in LREE to MREE whereas  
21  
22  
23 451 the HREE are almost flat ( $La_N/Nd_N = 2.63$ ;  $Sm_N/Ho_N = 2.07$ ;  $Er_N/Lu_N = 1.19$ ). This sample  
24  
25 452 shows a negative Eu anomaly ( $(Eu/Eu^*)_N = 0.85$ ).

26  
27 453 OCI rocks have complex REE patterns, showing varying degrees of enrichment in  
28  
29  
30 454 LREE ( $La_N/Nd_N = 1.76 - 3.80$ ) and relative depletion in MREE and HREE but with similar  
31  
32  
33 455  $Sm_N/Ho_N$  and  $Er_N/Lu_N$  ratios (0.35 and 0.82, respectively) compared to pure serpentinite  
34  
35 456 sample. Only sample ISC 13-1 shows a higher concentration of MREE ( $Sm_N/Ho_N = 0.89$ ). Of  
36  
37  
38 457 the OCI samples, ZRL 13-1 is the only rock with a positive Eu anomaly. The REE patterns of  
39  
40 458 the sedimentary ophicarbonates (OCII) are comparable to those of the OCI rocks, with the  
41  
42 459 exception of sample SG 15-2 that is far more enriched in all REE ( $La_N/Nd_N = 1.78$ ;  $Sm_N/Ho_N$   
43  
44  
45 460  $= 0.89$ ;  $Er_N/Lu_N = 0.93$ ), although the concentrations are all below those of primitive mantle  
46  
47 461 to which they are normalized. All ophicarbonates have negative Ce anomalies, with  $(Ce/Ce^*)_N$   
48  
49  
50 462 ranging from 0.12 to 0.80.

51  
52 463 Pumpellyite-actinolite-facies ophicarbonates (PL 13-1 and 2) are homogeneous in their  
53  
54  
55 464 abundances of REE with concentrations far lower than those of the primitive mantle. Overall,  
56  
57 465 their REE concentrations differ greatly from those of the oceanic ophicarbonates. More  
58  
59 466 specifically, they show enrichment in LREE ( $La_N/Nd_N = 2.77$ ) without negative Ce anomalies,  
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65

467 relatively flat MREE and HREE patterns ( $\text{Sm}_N/\text{Ho}_N = 1.09$ ;  $\text{Er}_N/\text{Lu}_N = 0.89$ ), and positive Eu  
1 468 anomalies ( $(\text{Eu}/\text{Eu}^*)_N = 1.81$ ).

2  
3 469 Blueschist-facies ophicarbonates CU3 and CU 13-1 have similar REE patterns,  
4  
5  
6 470 differing only in their HREE concentrations, the latter lower in sample CU 13-1 ( $\text{Er}_N/\text{Lu}_N =$   
7  
8 471  $2.50$  against  $1.05$ ). They display a slight enrichment in LREE ( $\text{La}_N/\text{Nd}_N = 0.74 - 0.48$ ), slight  
9  
10  
11 472 depletion in MREE ( $\text{Sm}_N/\text{Ho}_N = 1.25 - 1.35$ ) and positive Eu anomalies ( $(\text{Eu}/\text{Eu}^*)_N = 1.46 -$   
12  
13 473  $1.51$ ). For these samples, all PM-normalized REE concentrations fall between 0.1 and 1.

14  
15 474 The eclogite-facies ophicarbonate LP13 EOC1 shows LREE enrichment ( $\text{La}_N/\text{Nd}_N =$   
16  
17  
18 475  $2.24$ ) with a strong negative Ce anomaly, and enrichment of MREE relative to HREE  
19  
20  
21 476 ( $\text{Sm}_N/\text{Lu}_N = 0.54$ ). Europium shows a slight negative anomaly. Sample LP13 EOC3 shows a  
22  
23 477 similar REE pattern ( $\text{La}_N/\text{Nd}_N = 1.18$  and  $\text{Sm}_N/\text{Lu}_N = 0.62$ ), but about 1 order of magnitude  
24  
25 478 lower on the normalized plot. The negative Ce anomaly is less pronounced and Eu shows a  
26  
27  
28 479 slight positive anomaly. The eclogite-facies sample SA 15-1, from the S. Anna locality in the  
29  
30 480 Voltri Massif, displays a low  $\text{La}_N/\text{Nd}_N$  ratio ( $0.70$ ) and shows MREE enrichment relative to  
31  
32  
33 481 HREE ( $\text{Sm}_N/\text{Lu}_N = 3.64$ ). The magnesite-bearing ophicarbonate (SA 15-8) shows enriched  
34  
35 482 REE concentrations, relative to primitive mantle values, with LREE lower than MREE and  
36  
37  
38 483 HREE ( $\text{La}_N/\text{Lu}_N = 0.33$ ) and almost flat HREE ( $\text{Er}_N/\text{Lu}_N = 0.92$ ). Both SA 15-1 and SA 15-8  
39  
40 484 show negative Eu anomalies.

41  
42 485 The primitive mantle-normalized (PM, McDonough and Sun, 1995) trace elements  
43  
44  
45 486 concentrations of the serpentinite end-member (**Fig. 5C**) are similar to those of seafloor  
46  
47 487 serpentinites imparted by interaction with seawater-derived fluids. These similarities include  
48  
49  
50 488 enrichment in B in (up to 100 times PM) and positive anomalies in Li, Pb, As, Sb and Sr, but  
51  
52 489 always at concentrations lower than those of the primitive mantle. Cesium and Rb  
53  
54  
55 490 concentrations are below detection limits and Ba, U and Th concentrations also are very low.

56  
57 491 Carbonation processes lead to varying degrees of enrichment in LILE and fluid-mobile  
58  
59 492 elements (FME; e.g., As, Sb, U, Sr and Li) in the oceanic ophicarbonates OCI and OCII (**Fig.**  
60  
61  
62  
63  
64  
65

493 **5C)**. Similarity in the extents of enrichment in LILE, B, U, Th, Pb and Li is common for the  
1 494 subduction-related ophicarbonates (**Fig. 5D**). The trace element compositions of the eclogitic  
2  
3 495 samples are comparable to those of the other subduction-related ophicarbonates, with the  
4  
5  
6 496 main differences being their higher Pb, As and Sb concentrations. Niobium and Ta  
7  
8 497 concentrations in sample SA 15-8 are near primitive mantle values whereas the concentrations  
9  
10  
11 498 of these elements in all other samples are very low (between 0.002 and 0.1 normalized to PM  
12  
13 499 values).

14  
15 500

## 18 501 *5.2 Mineral chemistry and in situ trace element concentrations*

20 502 Major element compositions of representative analyses of minerals in each sample are  
21  
22  
23 503 reported in the **Table 2** and **3** and in the **Supplementary Table S1** and a detailed description  
24  
25 504 of the chemistry of the rock-forming minerals is reported in the **Supplementary Material 1**.  
26  
27  
28 505 Briefly, the serpentine from oceanic and subduction-related ophicarbonates shows major  
29  
30 506 element composition similar to that of serpentine in the pure serpentinite sample characterized  
31  
32  
33 507 by high MgO and SiO<sub>2</sub> contents and varying amounts of TiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub> and Cr<sub>2</sub>O<sub>3</sub> related to the  
34  
35 508 mantle mineral that was replaced (i.e., olivine or pyroxenes). Calcium-carbonate (CaCO<sub>3</sub>) is  
36  
37  
38 509 the dominant component of the carbonates in all oceanic and subduction-related  
39  
40 510 ophicarbonates (> 96%), except in eclogitic samples SA 15-8 and SA 15-1. The latter two  
41  
42 511 samples contain magnesite (MgCO<sub>3</sub> > 82% and FeCO<sub>3</sub> < 18%) and dolomite (with a Ca/Mg  
43  
44  
45 512 ratio near 1.2) as the most abundant carbonate minerals.

47 513 The *in-situ* trace element compositions of serpentines and carbonates from the studied  
48  
49  
50 514 samples are reported in **Table 2**, **Table 3**, and in **Figure 6**. When normalized to the primitive  
51  
52 515 mantle values (**Fig. 6A**), the trace element patterns of serpentine minerals from oceanic  
53  
54  
55 516 ophicarbonates show positive anomalies in Cs, U, B, As, Sb, Sr and Li, whereas other LILEs  
56  
57 517 and Th are depleted. Negative HFSE (High Field Strength Elements) anomalies also are  
58  
59 518 present. Compared to the trace element concentrations of the pure oceanic serpentines (grey  
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519 field in **Fig. 6A**), these serpentinites also are enriched in La with respect to Ce, compatible  
1 520 with the bulk-rock data for OCI and OCII. As expected, calcite in the OCI and OCII samples  
2  
3 521 (**Fig. 6C**) mostly shows enrichment in Sr with scattered enrichments in Rb, Ba, B, La and Li.  
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5  
6 522 Calcite from the less carbonated and less oxidized sample SG 15-4 has much higher U  
7  
8 523 concentrations than the calcite in the other samples. As for the serpentine minerals, La in  
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10  
11 524 calcite is enriched relative to Ce (e.g., sample MNT 15-2, **Fig. 6C**).

13 525 Antigorite from subducted ophicarbonates (**Fig. 6B**) shows trace element concentrations  
14  
15 526 comparable to those of the oceanic samples, except for variable Sr anomalies, with the  
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17  
18 527 concentrations in general lower than PM in the eclogitic samples and higher than PM in the  
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20  
21 528 all other rocks. Notably, As and Sb concentrations in the SA 15-1 and SA 15-8 samples fall at  
22  
23 529 the upper limit of the oceanic field. In situ trace element analyses of the calcites in sample PL  
24  
25 530 13-2 revealed concentrations comparable to those of oceanic OCI (**Fig. 6D**). In contrast, the  
26  
27  
28 531 calcites in the blueschist-facies ophicarbonates are depleted in Li and enriched in Pb and  
29  
30 532 MREE, relative to HREE, and show positive Eu anomalies (**Fig. 6D**) compared with calcite  
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32  
33 533 from the oceanic ophicarbonates. Carbonates in the eclogitic samples from La Pesca show  
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35 534 trace element patterns comparable to those of oceanic ophicarbonates; however, they show  
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37  
38 535 enrichment in W and Pb. Dolomite and calcite in sample SA 15-1 (S. Anna area) are  
39  
40 536 characterized by strong enrichment in LREE, Pb, Sr and MREE relative to HREE. Overall,  
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42 537 the concentrations of nearly all trace elements in calcite in this sample are one order of  
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44  
45 538 magnitude higher than those in the dolomites. Magnesite in sample SA 15-8 show low  
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47 539 concentrations in all trace elements, with very low Sr contents, but with strong enrichment in  
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49  
50 540 HREE relative to MREE and LREE.

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52 541

### 54 542 *5.3 Isotope geochemistry*

#### 57 543 *5.3.1 Carbon and oxygen isotope compositions of carbonates*

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544 Whole-rock C and O isotope compositions of the carbonate in the ophicarbonates suite  
1 545 analyzed here are listed in **Table 4**. The complete C and O isotope dataset for the micro-  
2  
3 546 drilled carbonates from each of the localities, itemized for texture and mineralogy, is provided  
4  
5  
6 547 in the **Supplementary Table S2** and the Figure in **Appendix A1**.

8 548 The  $\delta^{13}\text{C}$  of the oceanic carbonates ranges from about 0.0 to +3.0‰ (**Fig. 7A**), without  
9  
10 549 any significant difference between calcite replacing silicates or calcite in the veins (see the  
11  
12  
13 550 Figure in **Appendix A1**). These results are compatible with previously published data for the  
14  
15 551 Bracco ophicarbonates (Collins et al., 2015; Schwarzenbach et al., 2013). Carbonate in veins  
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17  
18 552 has  $\delta^{18}\text{O}$  values ranging from +18.2 to +26.0‰, whereas values for carbonates within the  
19  
20 553 serpentinite clasts and from the matrix in sedimentary rocks are lower, forming a cluster of  
21  
22  
23 554 +14.3 to 18.9‰ (**Fig. 7A**). The higher  $\delta^{18}\text{O}$  values correspond to calcite from late-stage veins  
24  
25 555 (**Fig. 2D**).

27 556 Pumpellyite-facies ophicarbonates (**Fig. 7B**) show  $\delta^{13}\text{C}$  within the range of the  
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29  
30 557 Apennine ophicarbonates (from 0.0 to +1.5‰) and the  $\delta^{18}\text{O}$  values form a tight cluster  
31  
32  
33 558 ranging from +14.4 to +15.5‰. The blueschist-facies CU3 samples show similar  $\delta^{13}\text{C}$  (+0.1  
34  
35 559 to +0.9‰) and  $\delta^{18}\text{O}$  from +15.8 and +17.9‰, but sample CU13-1 has negative  $\delta^{13}\text{C}$  value (-  
36  
37 560 2.7 to -1.1‰) and  $\delta^{18}\text{O}$  ranging from +16.4 to +18.1‰. The range shown by the latter sample  
38  
39  
40 561 is compatible with values obtained by Collins et al. (2015) for ophicarbonates and carbonated  
41  
42 562 metabasaltic rocks from the same area (**Fig. 7B**). Eclogitic sample LP13 EOC3 has  $\delta^{13}\text{C}$   
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44  
45 563 ranging from +0.7 to +1.4‰ and  $\delta^{18}\text{O}$  from 14.1 to 15.7‰, whereas LP13 EOC1 shows  
46  
47 564 similar  $\delta^{13}\text{C}$  (+0.8 to +1.6‰) but lower  $\delta^{18}\text{O}$  ranging from +10.8 and +11.5‰ (**Fig. 7B**).  
48  
49  
50 565 Other Voltri Massif ophicarbonates from the S. Anna locality are similar in their  $\delta^{13}\text{C}$  (+0.4 to  
51  
52 566 +1.7‰) but slightly higher in their  $\delta^{18}\text{O}$  (+11.6 to +12.7‰, **Fig. 7B**).

### 57 568 *5.3.2 Strontium isotope compositions*

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569 Strontium isotope compositions were analyzed for whole-rocks (except for sample SA  
1 570 15-1) and are listed in **Table 4** and are plotted in **Fig. 8**, against  $\delta^{18}\text{O}$  and  $\delta^{13}\text{C}$ . The  $^{87}\text{Sr}/^{86}\text{Sr}$   
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3 571 data for the Northern Apennines samples are reported for an initial time of 160 Ma, whereas  
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5  
6 572 the measured  $^{87}\text{Sr}/^{86}\text{Sr}$  ratio is reported for the subducted ophiicarbonates. This choice is based  
7  
8 573 on the fact that the extent of the Sr isotopic re-equilibration during subduction evolution is  
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10  
11 574 uncertain. However, as reported in **Table 4**, the differences between measured values and  
12  
13 575 age-corrected Sr isotope compositions for the subducted-related samples (for the Voltri  
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15 576 Massif of 33 Ma, Rubatto and Scambelluri, 2003) is very small (within the analytical error)  
16  
17  
18 577 due to the very low Rb/Sr of the samples.

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20 578 The  $^{87}\text{Sr}/^{86}\text{Sr}$  of the pure serpentinite end-member is 0.706929, which is compatible  
21  
22  
23 579 with values for serpentinite formed at the oceanic floor during interaction with seawater at  
24  
25 580 Jurassic time (Jurassic seawater value = 0.7070; Jones et al., 1994). The limestone is  
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28 581 characterized by higher Sr isotope value of 0.707048 which is compatible with a re-  
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30 582 equilibration with Upper Cretaceous seawater (Jones et al., 1994) as proposed by Barbieri et  
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32  
33 583 al. (1979). Ophiicarbonates of tectonic and sedimentary oceanic origins show  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios  
34  
35 584 similar to those of the pure serpentinite, ranging from 0.705746 to 0.706703. This range  
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38 585 confirms the Jurassic oceanic imprint of the ophiicarbonates due to interaction between  
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40 586 seawater-derived fluids in hydrothermal systems. The pumpellyite- and blueschist-facies  
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42 587 ophiicarbonates from Pietralavezzara and Ubaye show similar values ranging from 0.706383  
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45 588 to 0.706833. The only exception is for sample CU1 13-1 that displays a slight enrichment in  
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47 589 radiogenic Sr up to a value of 0.707530, well above the values for Jurassic and Cretaceous  
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49  
50 590 seawater. The  $^{87}\text{Sr}/^{86}\text{Sr}$  of sample LP13 EOC3 (0.706791) is comparable to that of oceanic  
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52 591 rocks, whereas the other eclogitic ophiicarbonates have far more radiogenic compositions:  
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54 592 sample LP13 EOC1 and SA 15-8 have  $^{87}\text{Sr}/^{86}\text{Sr} = 0.708453$  and  $0.709670$ , respectively, and  
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56  
57 593 calcite and dolomite separates from sample SA 15-1 have  $^{87}\text{Sr}/^{86}\text{Sr} = 0.707969$  and  $0.708130$ ,  
58  
59 594 respectively.  
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## 6. Discussion

Carbon release to subduction fluids and the mobility of aqueous-carbonic fluids within slabs or along the slab-mantle interface are actively debated topics in considerations of the solid Earth deep C cycle (Ague and Nicolescu, 2014; Berner, 1998; Collins et al., 2015; Cook-Kollars et al., 2014; Dasgupta and Hirschmann, 2010; Epstein et al., 2019; Jaeckel et al., 2018; Kelemen and Manning, 2015; Piccoli et al., 2016; Poli et al., 2009; Poli, 2015; Scambelluri et al., 2016). In this debate, a few studies have considered the geochemical evolution of ophicarbonates by investigating either the C-O isotope compositions of carbonates (Collins et al., 2015; Driesner, 1993; Lafay et al., 2017; Scambelluri et al., 2016; Weissert and Bernoulli, 1984), or the radiogenic isotope and trace element compositions of bulk-rocks and of rock-forming minerals (Miller and Cartwright, 2000; Yamaguchi et al., 2012). Lacking have been studies combining field and petrographic observation, and petrologic work, with trace element concentrations and C, O, and Sr isotope compositions of oceanic and of subduction-zone ophicarbonates aimed at elucidating the geochemical evolution of carbonated serpentinite along the subduction interface. Such studies can build on previous work employing trace elements and Sr isotopes to trace fluid-infiltration in subducting ultramafic rocks (Cannaò et al., 2016, 2015; Harvey et al., 2014). The approach presented in this paper employs such an integration and offers new insight regarding the mechanism of C mobilization in convergent margins.

### 6.1 Textural and petrologic features

The oceanic ophicarbonates studied here retain serpentine  $\pm$  magnetite mesh and bastite textures referred to as replacement of former mantle olivine and pyroxenes, respectively (from  $t_0$  to  $t_2$  in **Fig. 9**). During tectonic fragmentation of the serpentized mantle, carbonation of the bedrock by circulating hydrothermal fluids and accumulation of

621 serpentine clasts in sedimentary-tectonic breccias strongly affected the upper part of the  
1 622 ultramafic sequence (Schwarzenbach et al., 2013; Treves et al., 1995). Calcite formation after  
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3 623 the production of the mesh textures (**Fig. 2A**), and along the cleavage of pseudomorph  
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5 624 pyroxene, indicates that the mantle protolith was serpentinized prior to carbonation (t3 in **Fig.**  
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7  
8 625 **9**). Calcite formation after serpentinization requires open-system Ca and CO<sub>2</sub> input by the  
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11 626 hydrothermal fluids and the release of Si, Mg and water (as hydroxyl molecules – OH) into  
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13 627 the fluid. Thermodynamic modelling suggests that serpentine replacement by carbonates  
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15 628 occurred at 100 °C and 50 MPa at a high activity of CO<sub>2</sub> (Klein and Garrido, 2011). This  
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18 629 modelling was confirmed by experimental work showing that formation of calcite or  
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20 630 magnesite after serpentinization of peridotite is related to the activity of Mg<sup>2+</sup> in the fluid and  
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22  
23 631 to the presence of relict mantle minerals (Grozeva et al., 2017). Serpentine carbonation can  
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25 632 deliver abundant Mg and Si into the hydrothermal system with a consequent increase of the  
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27  
28 633 Mg/Ca and Si/Ca of the solutions. In the OCI and OCII ophicarbonates, the absence of Mg-  
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30 634 carbonates like dolomite and magnesite, and the only minor MgCO<sub>3</sub> component in the calcite,  
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32  
33 635 indicates that the hydrothermal fluids reached Mg<sup>2+</sup> saturation only after full consumption of  
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35 636 the aqueous CO<sub>2</sub>. In fact, the last carbonation event led to formation of late-stage calcite and  
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37  
38 637 talc-bearing veins (**Fig. 2D**). Talc crystallization in such late-stage veins reflects the chemical  
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40 638 evolution of hydrothermal fluids through time towards Mg- and Si-rich, and Ca- and CO<sub>2</sub>-  
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42 639 depleted composition, and marks the end of the oceanic hydrothermal carbonation system (t4  
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44  
45 640 in **Fig. 9**). **Figure 10** reports the results of thermodynamic calculations presented as an  
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47 641 activity-activity diagram (log(*a*SiO<sub>2</sub>) vs. log(*a*CO<sub>2</sub>)) simulating the oceanic carbonation of the  
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49  
50 642 pure serpentinite sample MNT 15-1 at low *T* and *P*. The initial carbonation occurs at  
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52 643 log(*a*CO<sub>2</sub>) values between -11.0 and -7.0 for a log(*a*SiO<sub>2</sub>) between -5.0 and -1.8 (red line in  
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54  
55 644 **Fig. 10**). In this ranges of log(*a*SiO<sub>2</sub>) values, the complete carbonation of relicts of residual  
56  
57 645 mantle minerals (e.g., olivine and pyroxenes) precedes the carbonation of serpentine, the latter  
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59 646 occurring at log(*a*CO<sub>2</sub>) of -6.0 to -5.0. This process is accompanied by release of H<sub>2</sub>O (blue  
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647 line in **Fig. 10**). At high  $\log(a\text{SiO}_2)$  values, talc becomes stable (above -1.8, see the green line  
1 648 in **Fig. 10**) over carbonates for a  $\log(a\text{CO}_2)$  greater than -6.5.

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3 649 The OCI and OCII ophicarbonates showing the highest modal proportions of rocks-  
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6 650 forming carbonate also display abundant hematite grains in the carbonated silicate domains  
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8 651 and hematite rinds on carbonated clasts (**Fig. 2A, B**; t4 in **Fig. 9**). Overall, these features  
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11 652 indicate highly oxidizing condition during carbonation, as also suggested by the formation of  
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13 653 hydrous andradite-rich garnet in the bastite textures after former pyroxenes (**Fig. 2C**).

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15 654 The above oceanic transformations created highly oxidized, hydrated, and carbonated  
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18 655 serpentinite reservoirs hosting up to 40wt%  $\text{CO}_2$  (**Fig. 4**). Subduction of these rocks was  
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20  
21 656 accompanied by deformation of the oceanic textures (Stöckhert, 2002) with displacement and  
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23 657 recrystallization of coarse oceanic calcite and serpentine along ductile shear zones (**Fig. 3**). At  
24  
25 658 this stage, deformation might have occurred under chemically closed- or open-system  
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27  
28 659 conditions. The behaviour of these rocks during subduction can be revealed by textural  
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30 660 analysis and study of geochemical markers. With the exception of samples SA 15-1 and SA  
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33 661 15-8, the majority of the subducted ophicarbonates for which data are presented here display  
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35 662 antigorite + calcite as the main mineral assemblage. Their textures and mineral assemblage  
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37  
38 663 could reflect an internal redistribution/recrystallization of minerals and elements inherited  
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40 664 from the oceanic stage. However, the pumpellyite- and the blueschist-facies samples are  
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42 665 characterized by growth of secondary calcite after the serpentinization of the porphyroclasts  
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44  
45 666 displaced along the shear-zone foliation (**Fig. 3B**). The geometry of these clasts suggests syn-  
46  
47 667 tectonic carbonation of serpentine during early subduction stages. Experimental studies  
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49  
50 668 showed that the solubility of calcite in  $\text{H}_2\text{O}$ -rich fluid increases with increasing  $P$ - $T$   
51  
52 669 conditions of the system (Caciagli and Manning, 2003), suggesting that the carbonation of the  
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54  
55 670 serpentinite clasts may have resulted from *in-situ* carbonate dissolution followed by *in-situ*  
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57 671 carbonate precipitation. Assuming a closed system condition, the  $\text{H}_2\text{O}$  required to trigger  
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59 672 calcite dissolution may have been derived from the serpentinite clast itself as few wt.% of  
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673 H<sub>2</sub>O is released during the serpentine phase transition at increasing *P-T* conditions (Schwartz  
1 674 et al., 2013).

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3 675 Additional textural evidence for C mobility during subduction is provided by the  
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5  
6 676 occurrence of carbonate veins crosscutting the eclogite-facies ophicarbonates in the Voltri  
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8 677 Massif (LP EOC13-1). Another indication of open system behaviour and infiltration of C-  
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10  
11 678 bearing fluids precipitating eclogite-facies magnesite veins is provided by the abundant fluid  
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13 679 inclusions hosted by porphyroclastic magnesite of sample SA 15-8. As shown in **Fig. 11D**,  
14  
15 680 and as discussed below, these inclusions shows significant enrichment in FME relative to the  
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17  
18 681 concentrations in their magnesite hosts. This might be further evidence for infiltration of C-  
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20 682 bearing fluids external to the ophicarbonate host.

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23 683 Taken together, the textures and other observations presented above, for distinct  
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25 684 domains in the ophicarbonates we investigated, provide a record of C mobilization over a  
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28 685 wide range of *P-T* conditions, beginning on the seafloor and during very early stages of  
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30 686 subduction.

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## 34 35 688 *6.2 Bulk and in-situ trace element features*

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37 689 It has been demonstrated that the serpentinization of the oceanic mantle rocks does not  
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40 690 strongly modify the REE composition of the peridotite protoliths (Deschamps et al., 2013 and  
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42 691 reference therein). The REE composition of the reference serpentinite MNT 15-1 thus  
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44  
45 692 reasonably reflects its mantle protolith that is comparable to Internal Liguride (Liguria, Italy)  
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47 693 and Monte Maggiore (Corsica, France) peridotites (grey field in **Fig. 5A**; Rampone et al.,  
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49  
50 694 2008, 1996). The similarity in the trace element composition of rock-forming serpentine from  
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52 695 MNT 15-1 (showing positive B anomaly and depletion in LILE and FME; grey field in **Fig.**  
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54 696 **6A**) with other oceanic serpentinites (e.g., Kodolányi et al., 2012) confirms that the alteration  
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57 697 process occurred in a marine environment. In contrast, the carbonation of oceanic  
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59 698 serpentinites significantly affects REE concentrations (**Fig. 5A**; Allen and Seyfried, 2005).

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699 Accordingly, compared to serpentinite MNT 15-1, all of our oceanic ophicarbonates display  
1 700 enrichments in La and variable depletion in Ce (**Fig. 5A**). The oxidation process that occurs  
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3 701 during carbonation, revealed by conversion of magnetite to hematite, also affects the  
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5 702 behaviour of Ce that, upon oxidized to  $Ce^{4+}$ , becomes incompatible in the serpentine and  
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7  
8 703 carbonate. Cerium depletion is recorded by the bulk-rock and the mineral *in-situ* analyses of  
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10  
11 704 most oceanic ophicarbonates (**Fig. 5A** and **6A**). The enrichment of Ba, B, Cs, U, Pb, As, Sb,  
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13 705 Sr, Li in the oceanic ophicarbonates and their rock-forming serpentine and carbonate can be  
14  
15 706 also linked to the carbonation and oxidation of these rocks (Andreani et al., 2014).  
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17  
18 707 Enrichment of FME in serpentine may be explained by the alkaline conditions of the  
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21 708 hydrothermal environment increasing adsorption of these elements onto the serpentine  
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23 709 mineral structure. According to the experimental work by Lafay et al., (2016) the As and Sb  
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25 710 enrichments in serpentine (**Fig. 6A**) can be explained when considering that  $As^{5+}$  and  $Sb^{5+}$  are  
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27  
28 711 favourably sequestered by serpentine minerals, compared with  $As^{3+}$  and  $Sb^{3+}$ , the latter which  
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30 712 are mobilized in the fluid phase and/or precipitate as Sb-Ni-accessory phases. For the  
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32  
33 713 ophicarbonate calcite, the *in-situ* trace element concentrations are similar to those of  
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35 714 serpentine, at least for As that is variably enriched. Uranium is mostly enriched in calcite in  
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37  
38 715 the less oxidized sample SG15-4, consistent with its high solubility at elevated oxygen  
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40 716 fugacity (e.g., Bailey and Ragnarsdottir, 1994; Keppler and Wyllie, 1991). On the contrary,  
41  
42 717 the Ce concentration in this sample is less depleted, perhaps indicating a key influence of  
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44  
45 718 oxygen fugacity in the Ce storage by ophicarbonate minerals.

47 719 The trace element patterns of the majority of the subducted ophicarbonates, and the  
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49  
50 720 minerals therein, do not display significant differences from the patterns in the oceanic  
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52 721 samples (**Figs. 5B, D** and **6B, D**). In particular, the coarse and recrystallized calcite in most  
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54 722 deformed subduction-related samples do not show significant differences in their trace  
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56  
57 723 element budgets (**Fig. 11A, B** and **C**). This evidence suggests that C-mobilization in these  
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59 724 samples occurred under closed-system conditions. The main differences in composition of the  
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725 subduction-related ophicarbonates from the oceanic ones are demonstrated by data for the  
1 726 eclogitic samples SA15-1 and SA 15-8, which show the highest REE concentrations (note the  
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3 727 enrichment in LREE relative to MREE and HREE in sample SA15-1, **Fig. 5B**) and higher Pb,  
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5 728 As, Sb compared with concentrations in oceanic precursors (**Fig. 5D**). These bulk-rock  
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7 729 anomalies are matched by enrichment in Cs, La, Pb, As, Sb of the rock-forming dolomite and  
8  
9 730 calcite (**Figs. 6B** and **D**). Enrichment in these elements in ultramafic rocks have been  
10  
11 731 attributed to exchange with sediment-derived fluids during subduction (Cannaò et al., 2016,  
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13 732 2015; Deschamps et al., 2011; Lafay et al., 2013). Therefore, the bulk-rock and *in-situ* trace  
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15 733 element data suggest that both closed and open system conditions were experienced by  
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17 734 ophicarbonates during the subduction, over their protracted metamorphic histories. In sample  
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19 735 SA 15-8, the vein magnesite traps primary and pseudo-secondary inclusions (**Fig. 3G, H**) of  
20  
21 736 the fluid that assisted magnesite growth. As expected, the LA-ICP-MS analyses of inclusion-  
22  
23 737 free and of inclusion-bearing magnesite (**Fig. 11D**) shows that the latter displays much higher  
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25 738 Cs, Ba, B, As, Sb, Rb, W, Pb concentrations. This indicates that the fluid carried crustally-  
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27 739 derived trace elements, either sourced in subducted sedimentary sequences, or released by  
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29 740 dehydration of previously metasomatized altered oceanic crust or associated serpentinites  
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31 741 (e.g., Cannaò and Malaspina, 2018 and reference therein).  
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### 42 743 *6.3 Tracing C mobility with 3D isotopes*

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44 744 The oceanic ophicarbonates studied here, like other oceanic ophicarbonates from the  
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46 745 Alps and Pyrenees (e.g., Clerc et al., 2014; Lafay et al., 2017), show wide ranges in O and C  
47  
48 746 isotope compositions (**Fig. 7A**), depending on the *T* of serpentinization and on the amount of  
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50 747 reduced C contributing to the  $\delta^{13}\text{C}$  of the fluids involved (e.g., Alt et al., 2012b; Delacour et  
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52 748 al., 2008). In the oceanic ophicarbonates presented here, the O and C isotope compositions of  
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54 749 calcite are comparable to published values for the Northern Apennine area (Collins et al.,  
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56 750 2015; Galli and Togliatti, 1965; Schwarzenbach et al., 2013). The calcite-seawater isotopic  
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751 equilibrium indicates that ophicarbonated calcite grew after serpentine over a  $T$  range of 130 –  
1 752 110 °C (assuming  $\delta^{18}\text{O}_{\text{seawater}} = 0\text{‰}$ , O'Neil et al., 1969) in OCI and OCII. Lower  $T$  of about  
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3 753 110 – 90 °C were obtained for calcite in veins cross-cutting all oceanic structures in OCI and  
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6 754 OCII. These  $T$  ranges are comparable to those calculated by Schwarzenbach et al. (2013) for  
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8 755 ophicarbonates from the same area (between 150 and 50 °C). All oceanic ophicarbonates  
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11 756 presented here preserve a C isotope fingerprint consistent with acquisition of C during  
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13 757 interaction with seawater-derived fluids. In OCII rocks, the matrix calcite in less oxidized  
14  
15 758 samples (SG 15-4 and SG 15-2) shows slightly lower  $\delta^{13}\text{C}$  likely as the result of interaction  
16  
17  
18 759 with serpentinite clasts, showing isotopically negative values (Alt et al., 2012b; Lafay et al.,  
19  
20 760 2017). The high degree of variability in the C-O isotope compositions of the oceanic protolith  
21  
22  
23 761 has to be taken into account when discussing the evolution of ophicarbonates during  
24  
25 762 subduction evolution. The  $^{87}\text{Sr}/^{86}\text{Sr}$  ratios of the oceanic ophicarbonates investigated here and  
26  
27  
28 763 of the pure serpentinite MNT 15-1 range between 0.7060 and 0.7070 (**Fig. 8**): such values  
29  
30 764 approach the value of 0.7070 defined for Jurassic seawater (Jones et al., 1994), consistent  
31  
32  
33 765 with carbonation having involved Jurassic seawater re-equilibrated with a mafic/ultramafic  
34  
35 766 reservoir.

37 767         Carbonate  $\delta^{18}\text{O}$  values for the pumpellyite- and of the blueschist-facies ophicarbonates  
38  
39  
40 768 overlap the lower- $\delta^{18}\text{O}$  range for OCI and OCII. The  $T$  range achieved for calcite  
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42 769 crystallization in these rocks, based on  $\delta^{18}\text{O}$  values, is 130-90 °C (assuming  $\delta^{18}\text{O}_{\text{seawater}} = 0\text{‰}$ ,  
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44  
45 770 O'Neil et al., 1969). Such  $T$  are in the range for alteration of the oceanic ophicarbonates,  
46  
47 771 indicating that these ophicarbonates retain the oceanic O isotope imprint of the protolith rocks.  
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49  
50 772 Therefore, although subducted ophicarbonates can display either undeformed or deformed  
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52 773 textures attributable to dynamic subduction-zone recrystallization, their  $\delta^{18}\text{O}$  does not require  
53  
54 774 contribution by alteration after seawater-rock interaction. Whereas the O isotope  
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56  
57 775 compositions of these rocks are consistent with little or no subduction-related modification of  
58  
59 776 seafloor compositions, the  $\delta^{13}\text{C}$  values (particularly that of sample CU 13-1) could reflect  
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777 some modification by high-*P* metamorphic fluids. The  $\delta^{13}\text{C}$  values of calcite in the  
1 778 pumpellyite-facies samples and the undeformed blueschist-facies sample CU3 show no  
2  
3 779 significant deviation from the values expected for oceanic ophicarbonates (**Fig. 7B**), despite  
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5  
6 780 the sampling of two generations of early (coarse) and later (recrystallized) calcite in these  
7  
8 781 subduction-zone metamorphosed samples (see **Appendix Figure A1**). Coupling this  
9  
10 782 observation with the preservation of the oceanic  $\delta^{18}\text{O}$ , with the  $^{87}\text{Sr}/^{86}\text{Sr}$  signature (**Fig. 8**),  
11  
12 783 and the *in-situ* trace element analyses (**Fig. 11A**) indicates inheritance of the oceanic  
13  
14 784 compositions, at closed-system conditions, during subduction. In contrast, calcite in the  
15  
16 785 deformed sample CU 13-1 has lower negative  $\delta^{13}\text{C}$  values (from -1 to -2‰, **Fig. 7B**) that can  
17  
18 786 be ascribed to either larger contribution from reduced/organic C during oceanic carbonation  
19  
20 787 or the influence of reduced C during interaction with fluids sourced during subduction from  
21  
22 788 the surrounding metasedimentary rocks (Collins et al., 2015; Cook-Kollars et al., 2014). The  
23  
24 789 more radiogenic Sr isotope composition of CU 13-1 (**Fig. 8**) supports the second scenario (a  
25  
26 790 metasedimentary source), despite the lack of evidence for significant trace element influx  
27  
28 791 from sediments based on the *in-situ* analyses (**Fig. 11B**). Carbon additions from  
29  
30 792 metasedimentary rocks into ultramafic ophicarbonates in the Queyras area has been  
31  
32 793 documented by means of Fe and Zn isotope systematics by Debret et al. (2018), who showed  
33  
34 794 that during subduction, these rocks underwent several stages of carbonate dissolution and  
35  
36 795 precipitation involving sediment-derived fluids.  
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45 796 One eclogitic sample from the Voltri Massif (LP13 EOC3) affected by ductile  
46  
47 797 deformation and carbonate recrystallization displays  $\delta^{13}\text{C}$  and  $\delta^{18}\text{O}$  overlapping values for the  
48  
49 798 pumpellyite-facies rocks (**Fig. 7B**), suggesting a subduction evolution under closed-system  
50  
51 799 conditions. This interpretation is supported by a lack of enrichment in FME of the  
52  
53 800 recrystallized calcite (**Fig. 11C**) and by low radiogenic Sr, with  $^{87}\text{Sr}/^{86}\text{Sr}$  still compatible with  
54  
55 801 an oceanic fingerprint (**Fig. 8**). In contrast, the other eclogitic samples from Voltri Massif are  
56  
57 802 characterized by the lowest  $\delta^{18}\text{O}$  (down to +10.5‰, **Fig. 7B**), matching with the compositions  
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803 of the eclogite-facies ophicarbonates from the Zermatt-Saas Zone (Collins et al., 2015),  
1 804 whereas their  $\delta^{13}\text{C}$  values are similar to those of the other oceanic and subducted  
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3 805 ophicarbonates (**Fig. 7B**). Considering that in these samples, (i) magnesite veins formed  
4  
5 806 during subduction (sample SA 15-8), (ii) dolomite crystallized together with antigorite and  
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8 807 diopside under high- $P$  conditions (sample SA 15-1), and (iii) calcite veins crosscut the  
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10  
11 808 prograde antigorite (sample LP13 EOC1), and assuming that their original  $\delta^{18}\text{O}$  was the same  
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13 809 as that of the oceanic ophicarbonates, the Voltri ophicarbonates are regarded as having re-  
14  
15 810 equilibrated with metamorphic fluids that lowered their seafloor-inherited  $\delta^{18}\text{O}$  (e.g., Miller  
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17  
18 811 and Cartwright, 2000). Decrease in the  $\delta^{18}\text{O}$  of these samples is coupled with increase in the  
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20 812 FME budgets of carbonates (**Fig. 6**) and with the bulk-rock enrichment in radiogenic Sr (**Fig.**  
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22 813 **8**), indicating influx of externally-derived fluids.

25 814         Carbonic fluids sourced from metasedimentary rocks during prograde subduction  
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27 815 should be characterized by high  $\delta^{18}\text{O}$  and low  $\delta^{13}\text{C}$ , likely with radiogenic  $^{87}\text{Sr}/^{86}\text{Sr}$  reflecting  
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29  
30 816 the elevated Rb/Sr of the rocks (Cook-kollars et al., 2014; Yamaguchi et al., 2012). The  $\delta^{18}\text{O}$   
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32 817 values of calcite, dolomite and magnesite in the eclogitic ophicarbonates suggest that the  
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34  
35 818 fluids in equilibrium with these minerals should have  $\delta^{18}\text{O}$  values of +6.7 to +10.5‰ (at peak  
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37 819  $T$  of 500 °C), i.e., +3.8 to +2.5‰ lower than the associated minerals (using the  $\Delta^{18}\text{O}_{\text{carbonate-}}$   
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40 820  $\text{water}$  of Zheng, 1999). Such isotopic values are compatible with derivation of the fluids from  
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42 821 altered oceanic crust and/or serpentinitized mantle (Cartwright and Barnicoat, 1999) rather than  
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45 822 sedimentary sources (see Jaekel et al., 2018). However, the Sr isotope ratios of oceanic,  
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47 823 mafic crustal reservoirs are generally less radiogenic than values for the ophicarbonates  
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49  
50 824 analyzed in this study. The conundrum is possibly solved by associating the isotopic imprints  
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52 825 observed in the high- $P$  ophicarbonates with interaction with fluids released by devolatilization  
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54 826 of hybridized ultramafic rocks bearing some sedimentary signatures. As a related example,  
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57 827 Cannà et al. (2016) demonstrated that interaction of sediment-derived fluids with the Voltri  
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59 828 serpentinites caused enrichments in FME and provided radiogenic Sr and Pb. The subduction  
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829 dehydration of such a hybrid serpentinite could have released hydrous fluids contributing the  
1 830 isotopic signatures to the subduction ophicarbonates analyzed in this study (high  $^{87}\text{Sr}/^{86}\text{Sr}$ ,  
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3 831 low  $\delta^{18}\text{O}$ , FME enrichment). Beginning with serpentinites with + 6 / +8‰  $\delta^{18}\text{O}$  (Cannaò et al.,  
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6 832 2016; Früh-Green et al., 2001), the fluids released at 400 to 650 °C, corresponding to the  
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8 833 temperatures of olivine-in and antigorite-out reactions (Scambelluri et al., 1995), would have  
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10  
11 834  $\delta^{18}\text{O}$  ranging from +7.2 to +10.8‰. These values are in good agreement with those of the  
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13 835 metamorphic fluids in equilibrium with carbonates from the subduction-zone ophicarbonates  
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15 836 for which data are presented in this paper. The fluid generated by devolatilization of the  
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18 837 serpentinite also would have low C/O ratio (Alt et al., 2012a), enabling retention of the C  
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21 838 isotope compositions of the carbonates, even at fairly high fluid-rock ratios, but modifying the  
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23 839  $\delta^{18}\text{O}$  towards lower values. Alternatively, interaction of fluids derived by de-serpentinization  
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25 840 with metasedimentary rocks along their flow paths in the subduction zone, could have led to  
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28 841 the hybridized fluid compositions. Other recent studies of extensively veined mafic rocks and  
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30 842 Schistes Lustrés adjacent to fault systems though to represent transient subduction interfaces  
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32  
33 843 has demonstrated shifts in  $\delta^{18}\text{O}$  to similar values and invoked infiltration by H<sub>2</sub>O-rich fluids  
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35 844 generated by dehydration of mafic and/or ultramafic rocks at greater depths (Angiboust et al.,  
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38 845 2017; Epstein et al., 2019; Jaekel et al., 2018).

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## 41 42 847 **7. Conclusions**

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45 848 In this study, we provide petrographic and geochemical constraints regarding fluid-  
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47 849 driven C mobility in ophicarbonates representing shallow levels of subduction zones and their  
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50 850 equivalents subducted to depths experiencing eclogitic *P-T* conditions. We characterize the  
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52 851 sources of these fluids, identifying cases where compositions can be explained by closed-  
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55 852 system conditions and others where open-system fluid-rock interaction appears to have been  
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57 853 required. We demonstrate that decarbonation and dissolution/precipitation processes  
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60 854 operating in ancient subduction zones are best traced using a combination of detailed field and  
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855 petrographic observations, FME inventories and C-O and Sr isotope systematics. We suggest  
1 856 that one challenge to our better understanding subduction zone C fluxes is the upscaling of  
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3 857 observations from field and geochemical studies of high-*P* metamorphic rocks to  
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5  
6 858 consideration of C cycling at Earth's modern consuming margins.  
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34  
35 870 manuscript and have been much appreciated.  
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## 40 872 **Figure Caption**

### 42 873 **Figure 1**

45 874 Simplified geological map of the Western Alps and location of the studied areas (modify after  
46  
47 875 Collins et al., 2015). (A) Geological map of Internal Liguride from the Ligurian N. Apennine  
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49 876 ophiolite (modified after Treves and Harper, 1994); (B) area of study from the Queyras/Ubaye  
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52 877 zone (modified after Michard et al., 2004); (C) geological map of the Voltri Massif (modified  
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54 878 after Cannà et al., 2016 and reference therein). Locations of the collected samples for all  
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57 879 three areas of study together with the respective sample' labels are indicated with the star  
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59 880 symbol.  
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**Figure 2**

1 882 **Figure 2**  
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3 883 Photomicrographs of the tectonic/hydrothermal (OCI) and sedimentary (OCII) ophicarbonates  
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6 884 from the Northern Apennines: (A) parallel and (B) crossed polarized light microphotographs  
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8 885 of completely carbonated serpentinite clast. Equant calcite grains overgrowing  
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11 886 pseudomorphic serpentine minerals in the cell of mesh texture and a film of hematite grown  
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13 887 along the border, giving them a ghost-like texture (OCI – MNT 13-5). (C) Former  
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15 888 porphyroclast of serpentinitized pyroxene (i.e., bastite; red dashed line) partially carbonated  
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18 889 coupled with formation of oceanic garnet along the cleavage (OCI – ISC 13-1). (D) Talc  
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20 890 aggregates overgrowing coarse grains of calcite in tardive veins. This represents the last  
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22  
23 891 process affecting the oceanic ophicarbonates (OCI – MNT 13-5). (E) Boytroidal versus  
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25 892 irregular blocky calcite in vein (OCII – SG15-4). (F) Calcite in sedimentary matrix rich in  
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27  
28 893 hematite. Several serpentinite (Serp) clasts with irregular shape are visible (OCII – MNT 15-  
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30 894 2).

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33 895**Figure 3**

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35 896 **Figure 3**  
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37 897 Photomicrographs of the subducted ophicarbonates. (A) Old elongated and strained calcite  
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39  
40 898 grains, partially recrystallized along grain boundaries. Complete recrystallization of un-  
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42 899 twinned calcites along major shear zone is visible (PL 13-2). (B) Re-oriented serpentinite  
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45 900 clast along the shear zone partially carbonated (Carb clast). Relict of twinned calcites are  
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47 901 visible enclosed in recrystallized prograde calcite mylonites (PL 13-2). (C) Stretched  
48  
49  
50 902 serpentinite clast (Serp) with microfractures and pressure shadows filled by syn-tectonic  
51  
52 903 extensional calcite-fibers (Syn-tect. calcite). Syn-tectonic calcite is probably linked to the  
53  
54 904 oceanic stage (CU3). (D) Dynamic recrystallization of old strained calcite porphyroclasts  
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56  
57 905 producing fine untwinned calcite grains (CU 13-1). (E) Green garnet (Andradite) associated  
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59 906 with oxide (Cr-magnetite – Cr-mgt) and calcite along major deformation zone (LP13 EOC1).  
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907 (F) Relict of the high pressure layer with dolomite + diopside + antigorite (SA 15-1). (G-H)

1 908 Primary and pseudo-secondary fluid inclusions (FI) hosted in magnesite (Mgs) from sample

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3 909 SA 15-8.

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7 911 **Figure 4**

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9 912 Bulk-rock major element data showing the relationship between CaO, MgO and SiO<sub>2</sub> of  
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13 913 ophicarbonates as well as variations in CO<sub>2</sub> and H<sub>2</sub>O with respect to MgO and CaO.

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15 914 Oceanic and subducted ophicarbonates fall in between pure serpentinite and limestone end-  
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17  
18 915 members.

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20 916

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22 917 **Figure 5**

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24  
25 918 Primitive mantle normalized (McDonough and Sun, 1995) Rare Earth Element (REE) and  
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27  
28 919 trace element patterns of ophicarbonates from oceanic (A, C) and subduction (B, D) stages. In

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30 920 (A) the grey field report the range of the peridotites from Internal Liguride (Rampone et al.,

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32 921 1996). In (B) and (D), the grey fields represent the REE and trace element pattern of the  
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34  
35 922 oceanic oph carbonate illustrated in (A) and (C).

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37 923

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39 924 **Figure 6**

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42 925 Primitive mantle normalized trace elements patterns of serpentine and carbonate from oceanic  
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45 926 (A, C) and subduction (B, D) stages. In (B) and (D) grey fields represent the trace element

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47 927 variation of the serpentine and calcite from oceanic origin illustrated in (A) and (C). Primitive  
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49  
50 928 mantle normalization from McDonough and Sun (1995).

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54 930 **Figure 7**

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57 931 Oxygen ( $\delta^{18}\text{O}_{\text{VSMOW}}$ ) and C ( $\delta^{13}\text{C}_{\text{VPDB}}$ ) isotopic signatures of ophicarbonates from oceanic  
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59 932 (A) and subduction (B) stages. In (A) oph carbonate fields from Bracco, the Pyrenean and

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933 Chenaillet (with different grey colours) are shown for comparison (from Collins et al., 2015;  
1 934 Clerc et al., 2014; Lafay et al., 2017). In **(B)** the Queyras and Zermatt-Saas fields are from  
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3 935 Collins et al., (2015) whereas Chatillon ophicarbonates field is from Driesner (1993). The  
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6 936 complete  $\delta^{18}\text{O}$  vs.  $\delta^{13}\text{C}$  dataset is shown in the **Appendix Figure A1** and in the  
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8 937 **Supplementary Table S2.**

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10 938

### 13 939 **Figure 8**

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15 940  $^{87}\text{Sr}/^{86}\text{Sr}$  vs.  $\delta^{18}\text{O}_{\text{VSMOW}}$  **(A)** and  $\delta^{13}\text{C}_{\text{VPDB}}$  **(B)** of ophicarbonates studied in this work. For  
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18 941 comparison the values of the pure oceanic serpentinite and limestone from the Northern  
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20 942 Apennines and of the Jurassic seawater are shown (dashed black lines).

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23 943

### 25 944 **Figure 9**

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27 945 Cartoon illustrating the sequence of hydration and carbonation/oxidation processes affecting  
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30 946 mantle rocks at oceanic conditions. See text for discussion.

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33 947

### 35 948 **Figure 10**

36  
37 949 Activity-activity diagram for Si and  $\text{CO}_2$  ( $\log(a\text{SiO}_2)$  and  $\log(a\text{CO}_2)$ , respectively) depicting  
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40 950 phase equilibria in the  $\text{MgO-CaO-Al}_2\text{O}_3\text{-SiO}_2\text{-H}_2\text{O-CO}_2$  system at 100 °C and 1500 bar  
41  
42 951 (calculated using the *Perple\_X* software, Connolly (2005) and the internal consistent  
43  
44  
45 952 thermodynamic database of Holland and Powell (1998) and the compensated Redlich-Kwong  
46  
47 953 (CORK) equation of state of Holland and Powell (1991) for  $\text{H}_2\text{O}$  and  $\text{CO}_2$  fluids). Key  
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49 954 carbonate-in (in red), talc-in (in green) and release  $\text{H}_2\text{O}$  (in blue) boundaries are emphasized.  
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52 955 See text for discussion.

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### 57 957 **Figure 11**

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958 Comparison between in situ trace elements data for coarse and recrystallized calcite from  
1 959 pumpellyite- (A), blueschist- (B) and eclogite-facies (C) samples. Trace element  
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3 960 compositions of fluid inclusion – magnesite mixed measurement compared with pure  
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6 961 magnesite are reported in (D). Primitive mantle data from McDonough and Sun (1995).  
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8 962  
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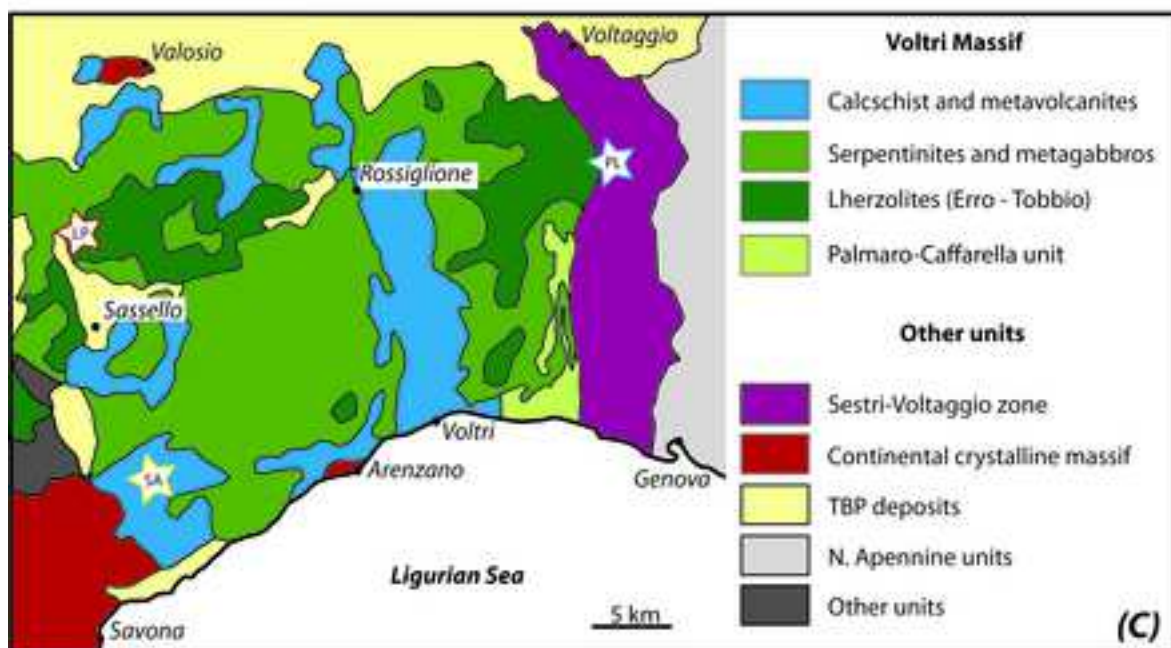
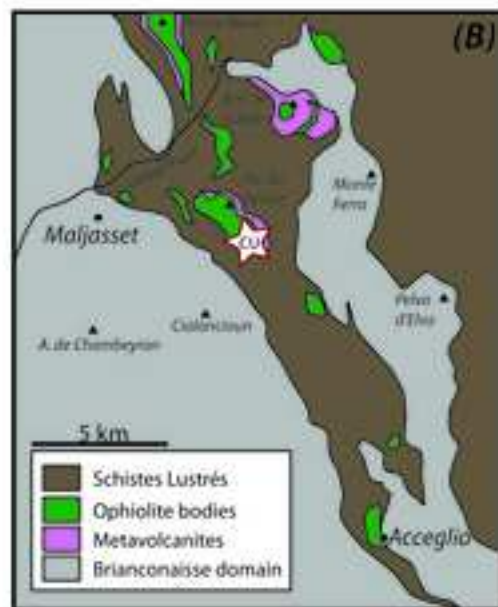
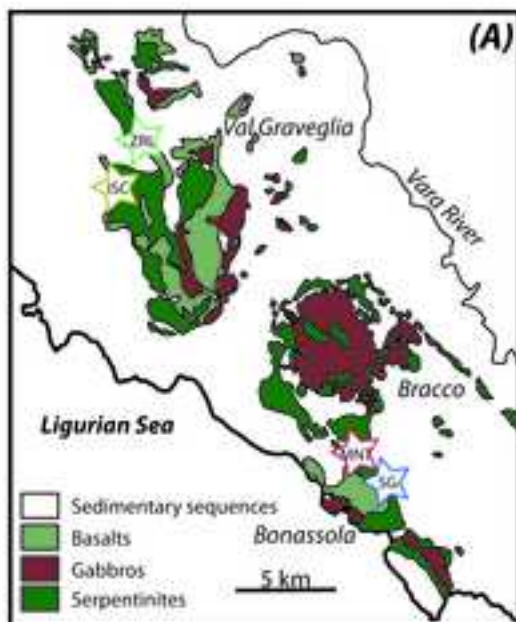


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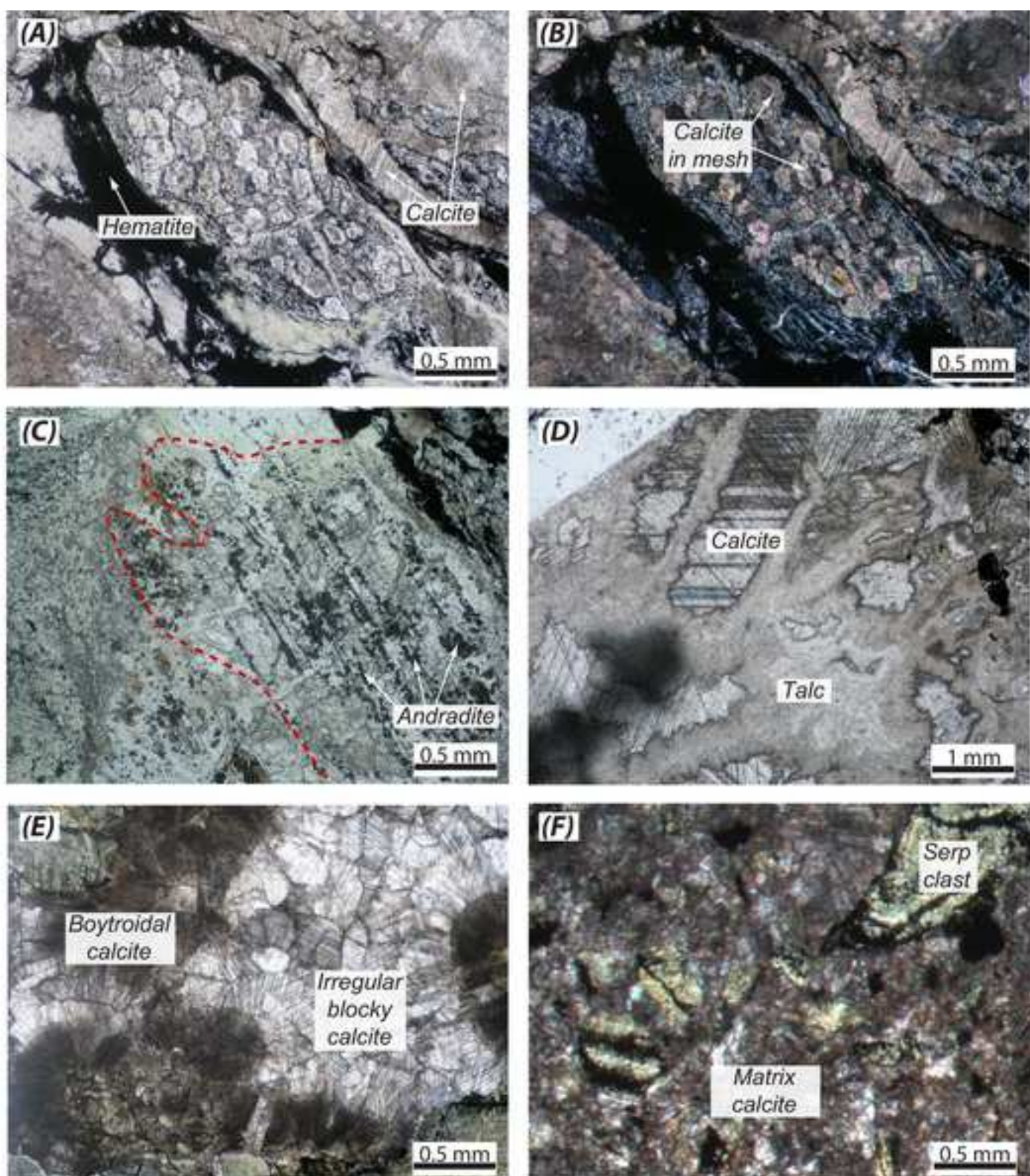


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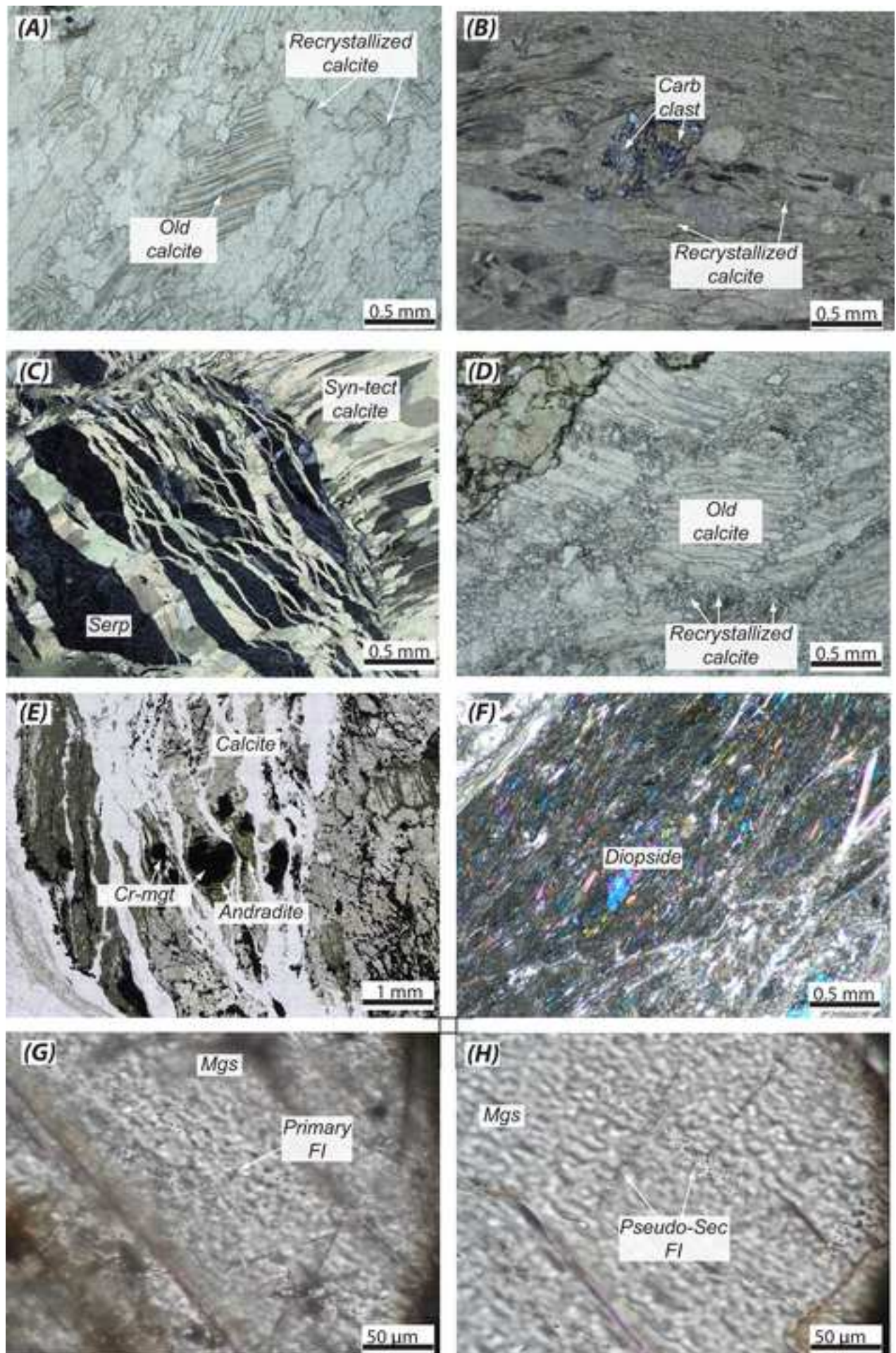


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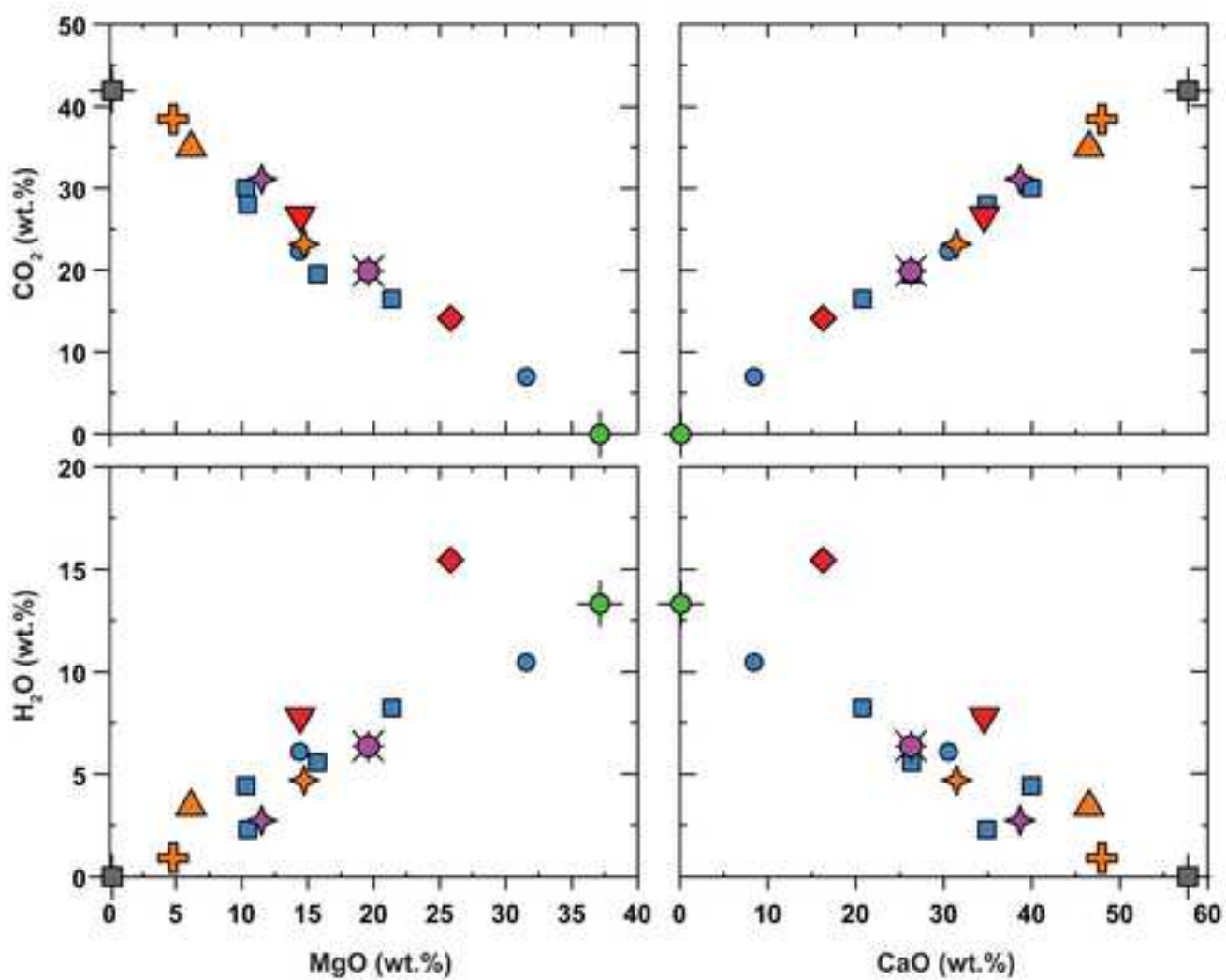
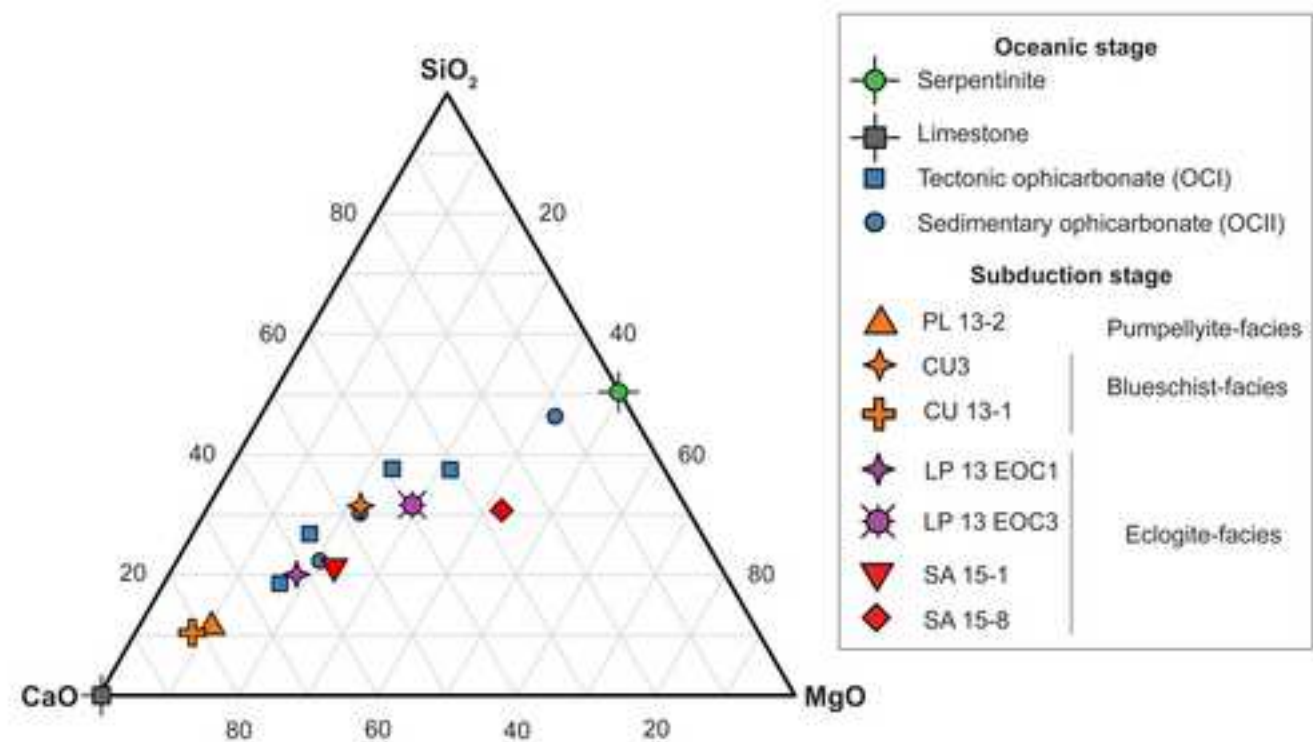


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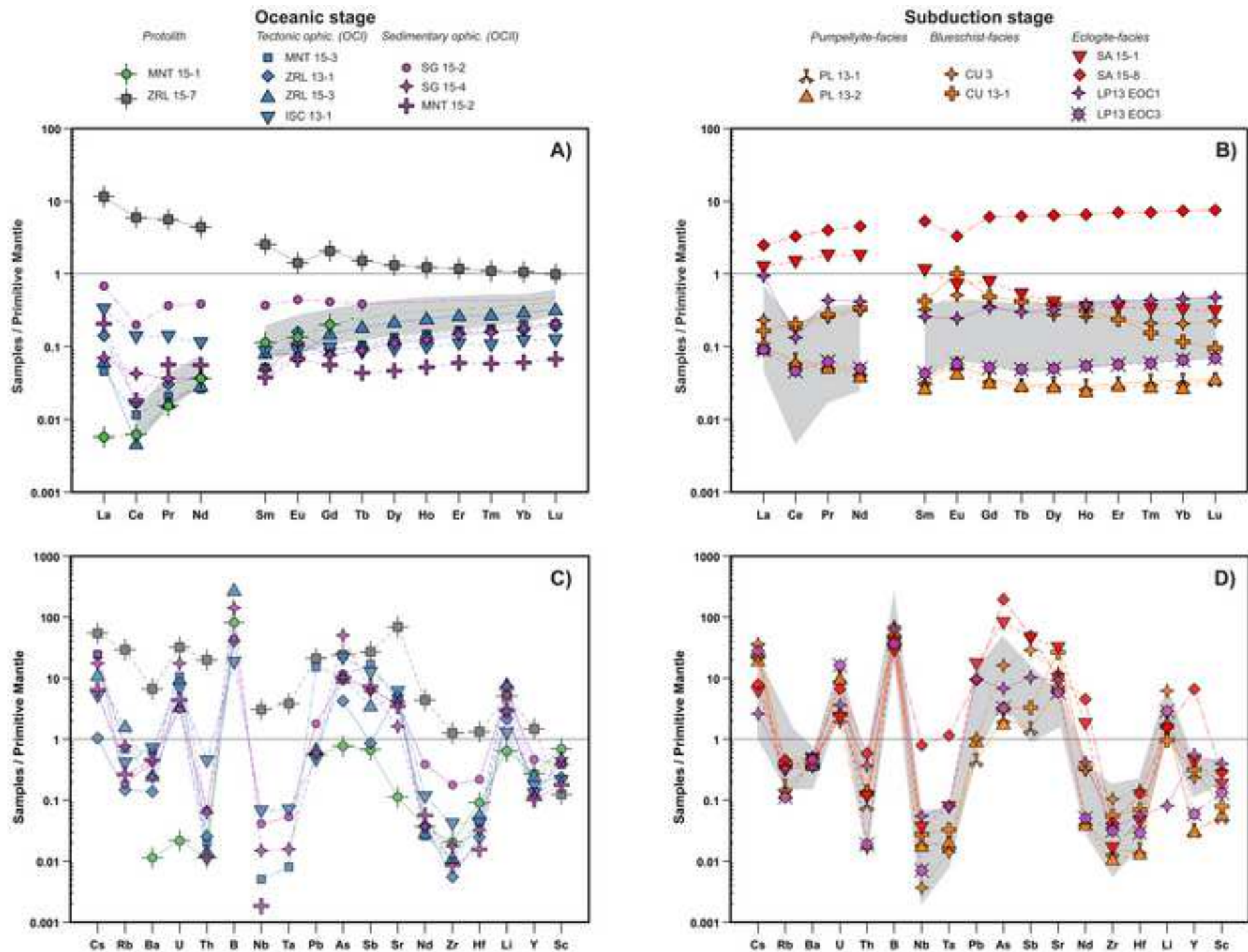


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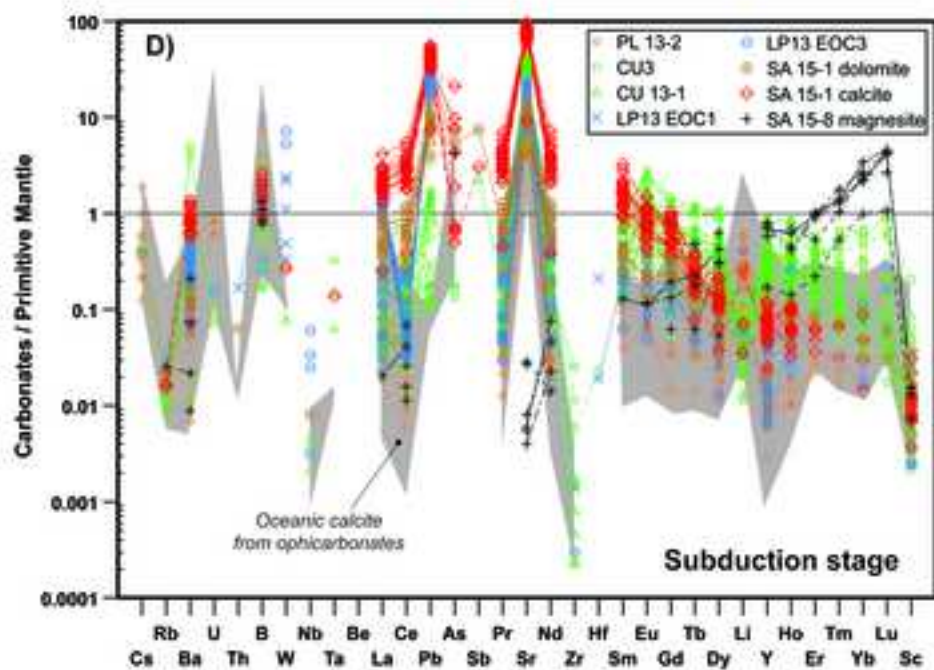
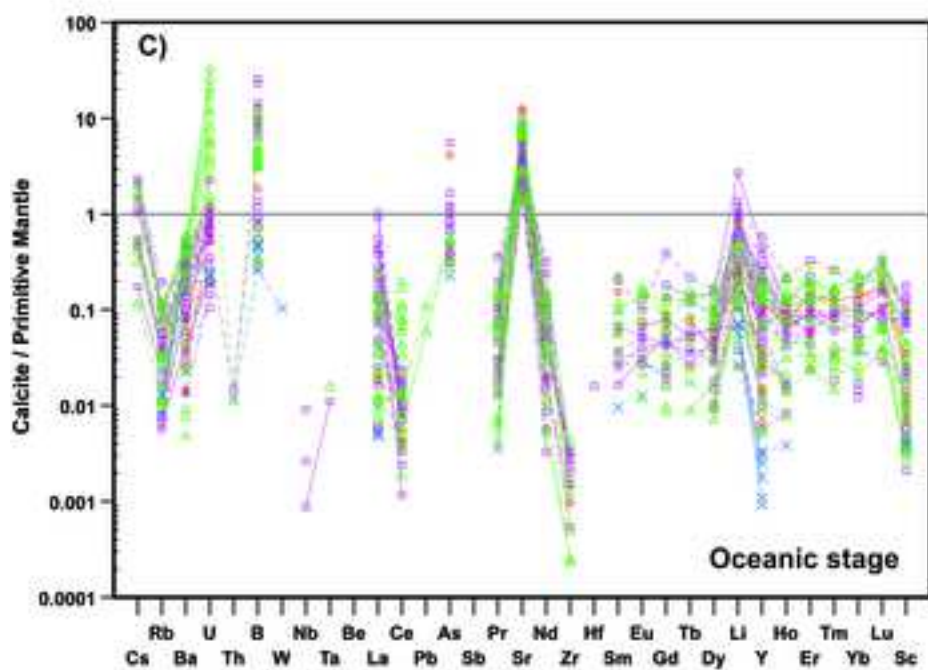
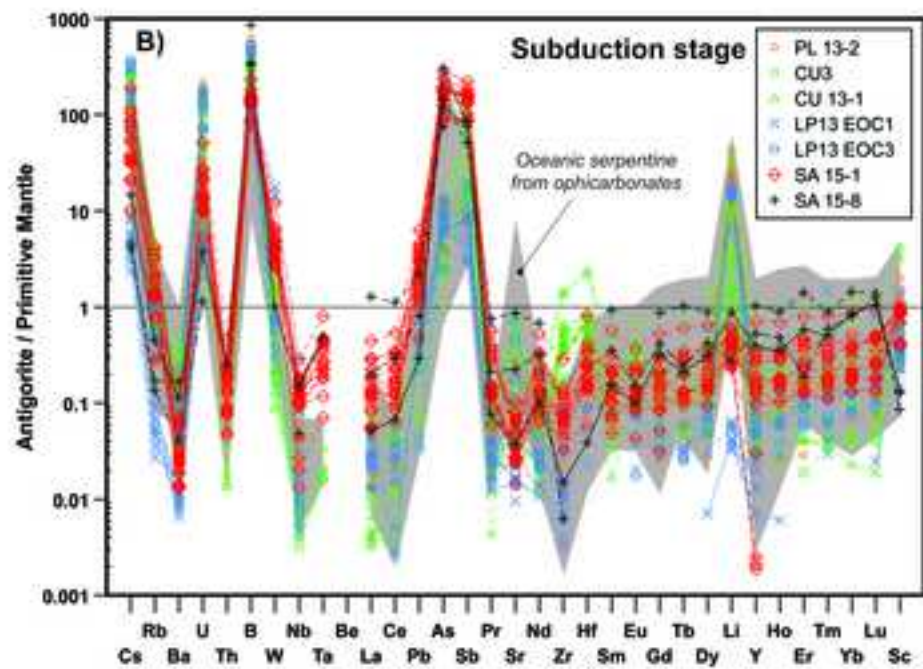
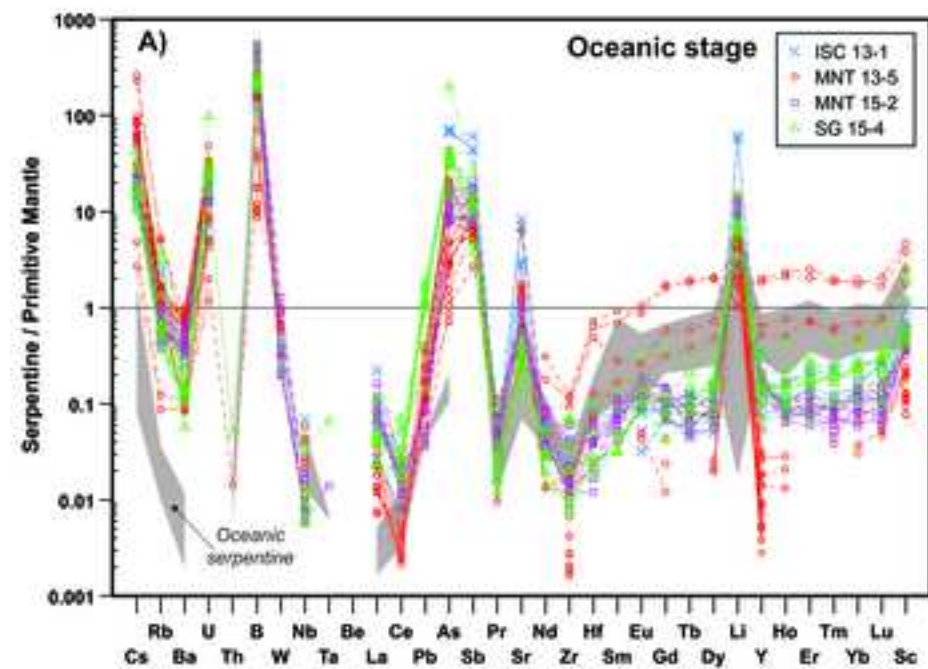


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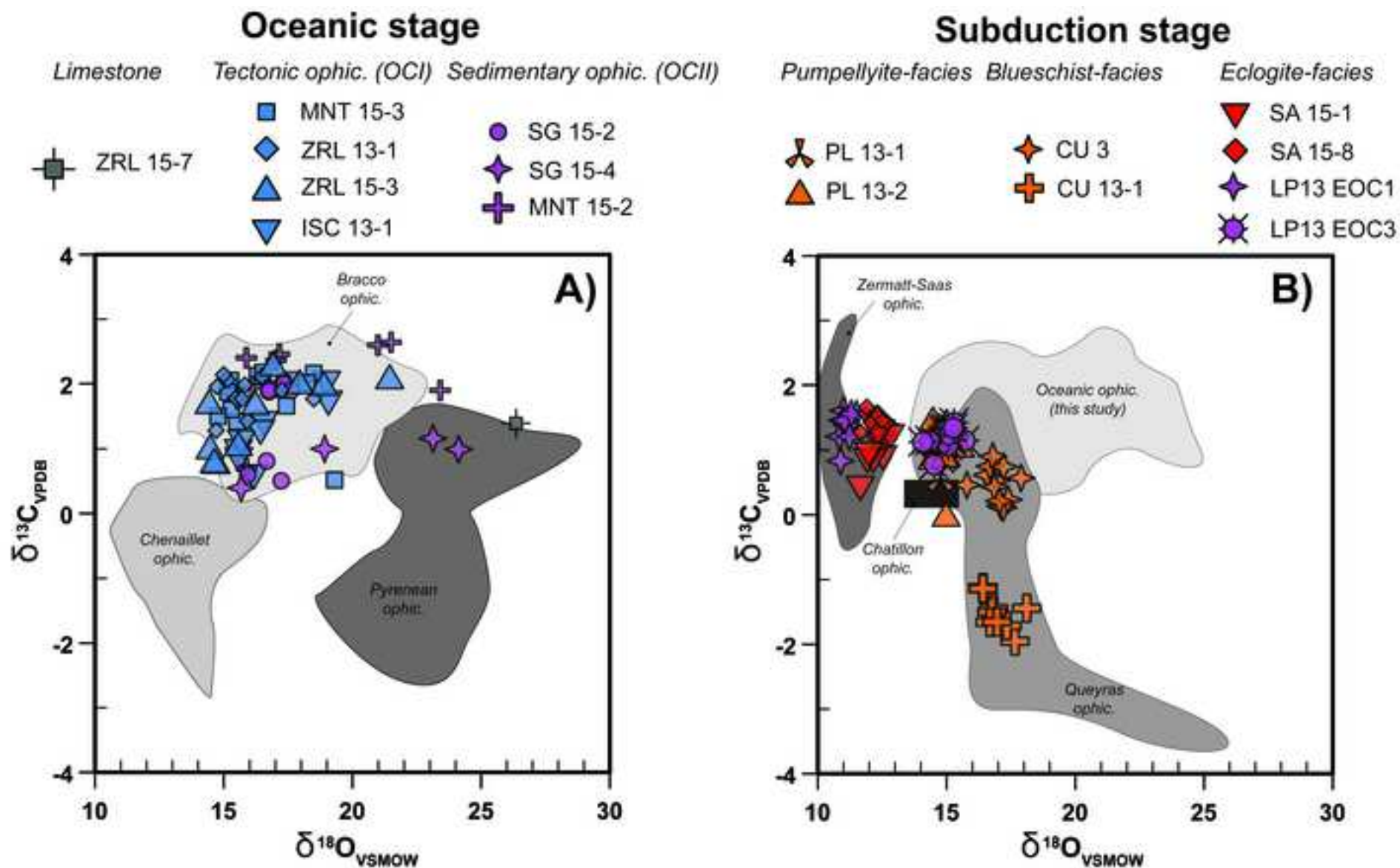


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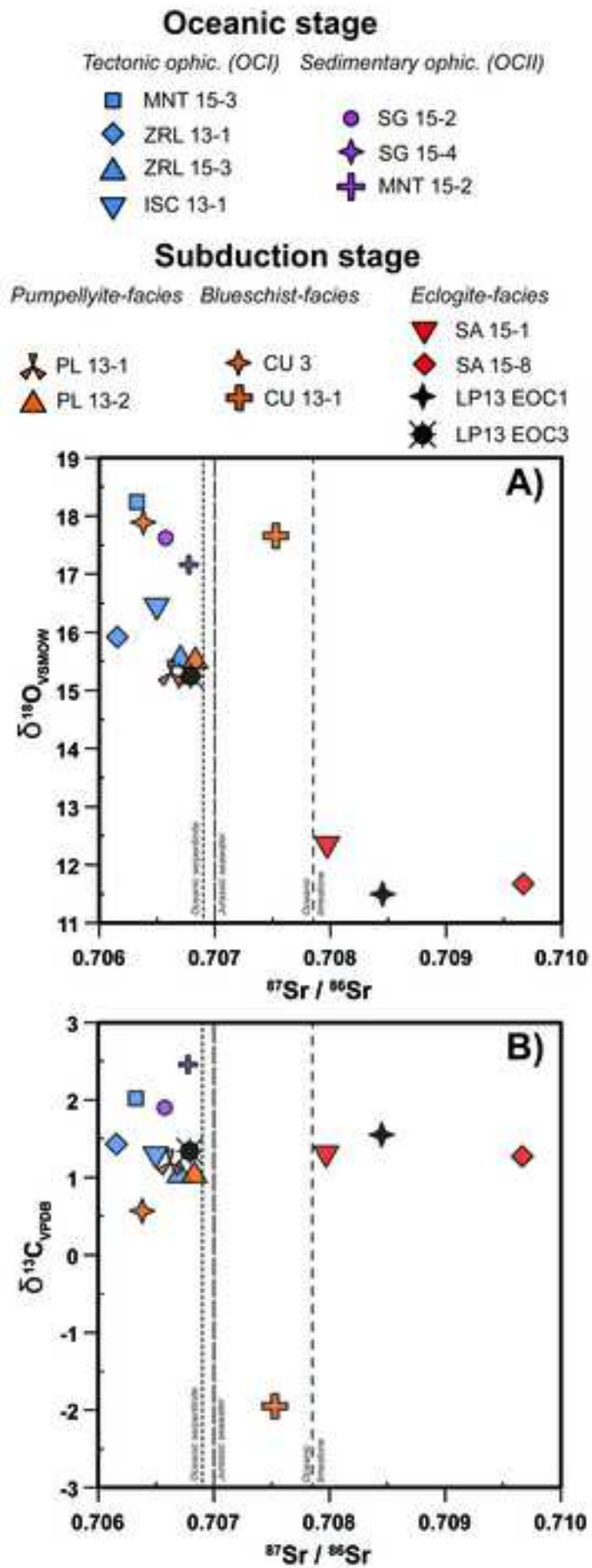


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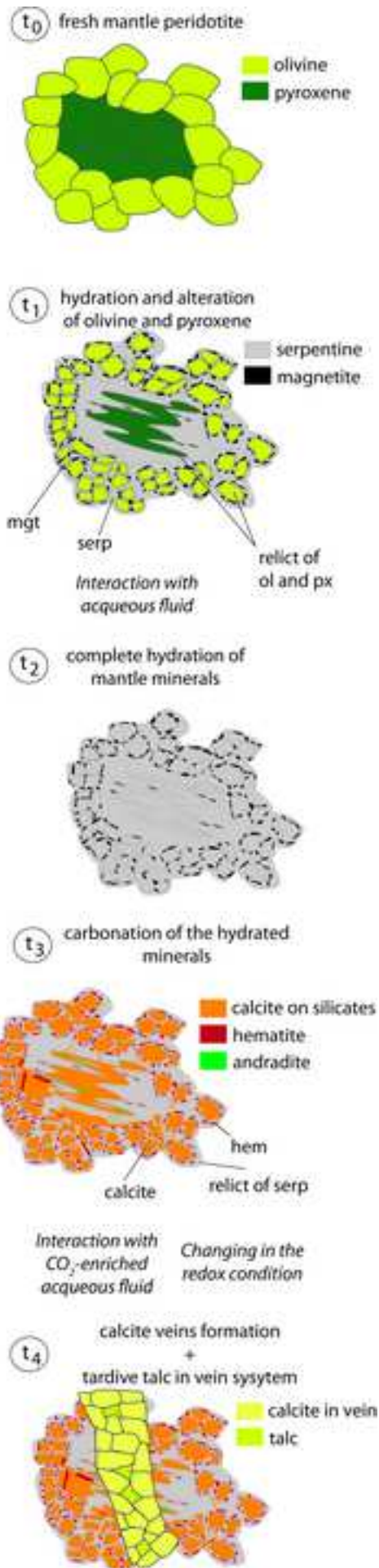


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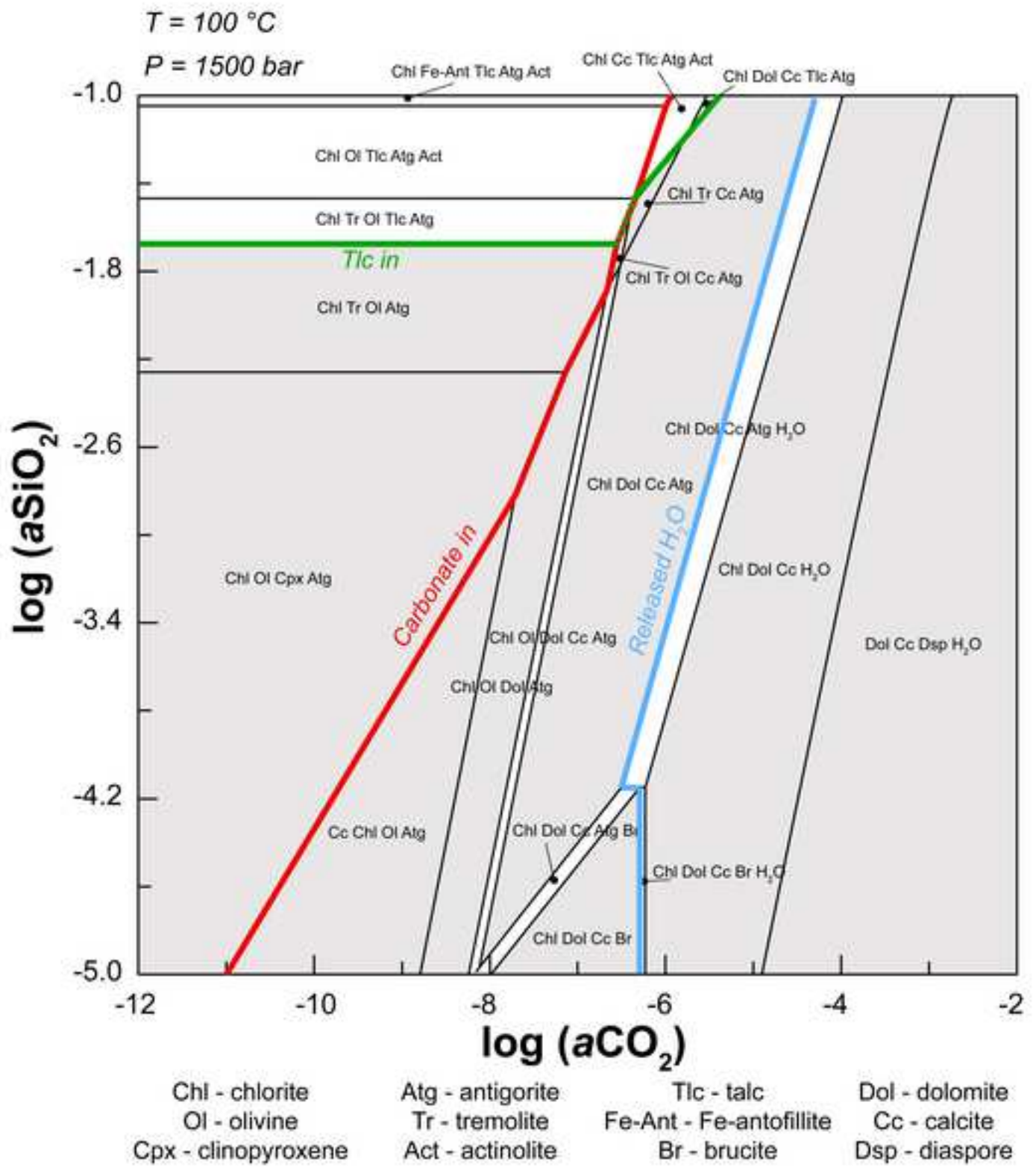
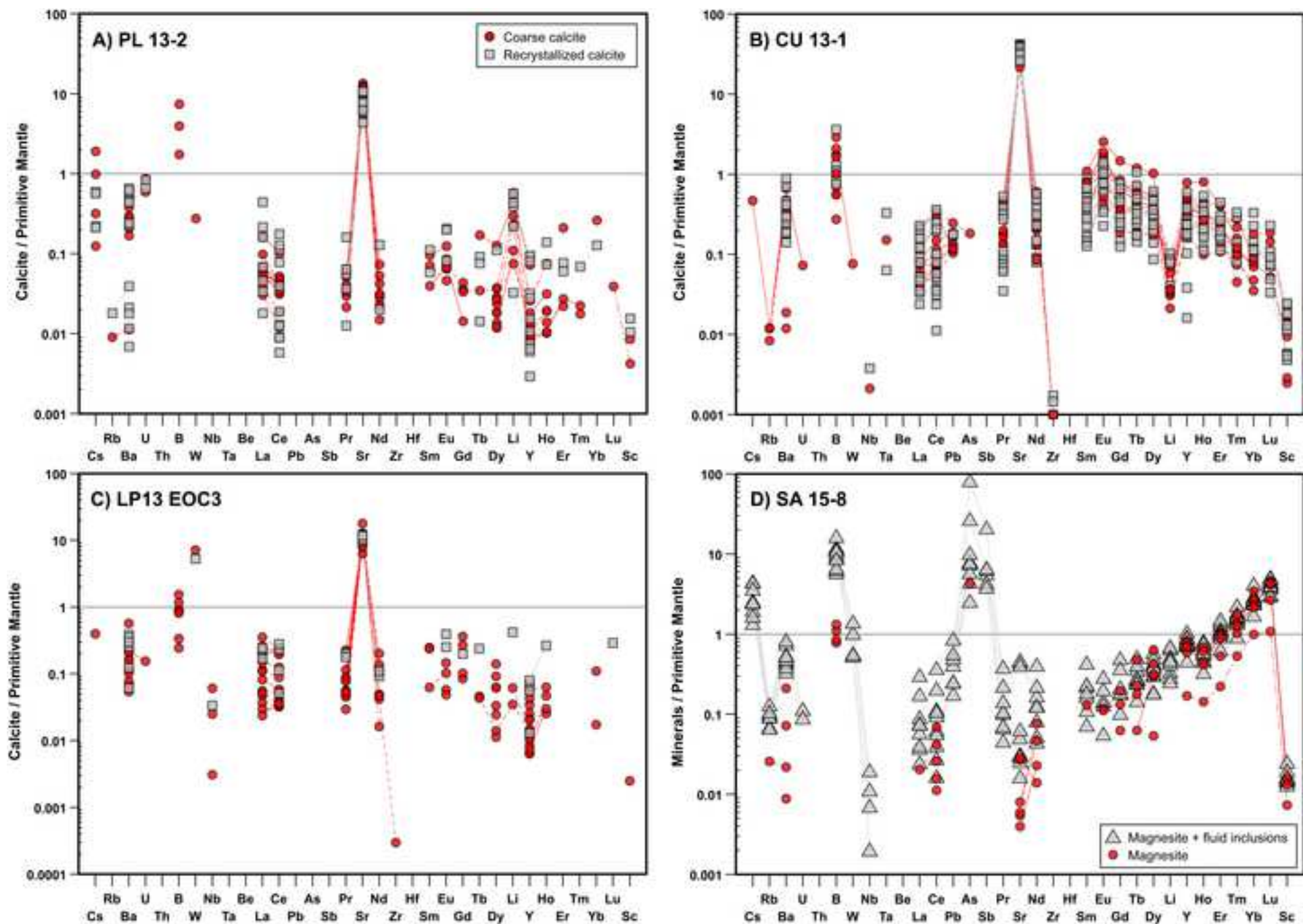


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**Table 1**  
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**Table 1.** Whole-rock major (wt.%) and trace element (ppm) analyses for serpentinite, limestone and ophicarbonate rocks

| GPS sample location                  | Oc. serp.               |                         | Limestone               |                         |                         |                         | Oc. tectonic ophicarbonates |                         |                         |                        | Oc. Sedimentary ophicarbonates |                         |                         |                         | Prograde ophicarbonates |                         |                         |        | Eclogite facies ophicarbonates |  |  |  |
|--------------------------------------|-------------------------|-------------------------|-------------------------|-------------------------|-------------------------|-------------------------|-----------------------------|-------------------------|-------------------------|------------------------|--------------------------------|-------------------------|-------------------------|-------------------------|-------------------------|-------------------------|-------------------------|--------|--------------------------------|--|--|--|
|                                      | MNT 15-1                | ZRL 15-7                | MNT 13-5                | ZRL 13-1                | ZRL 15-3                | ISC 13-1                | SG 15-4                     | SG 15-2                 | MNT 15-2                | PL 13-1                | PL 13-2                        | CU 3                    | CU 13-1                 | SA 15-1                 | SA 15-8                 | LP13 EOC1               | LP13 EOC3               |        |                                |  |  |  |
|                                      | N 44.19811<br>E 9.58860 | N 44.34558<br>E 9.45221 | N 44.19811<br>E 9.58860 | N 44.34558<br>E 9.45221 | N 44.34558<br>E 9.45221 | N 44.32808<br>E 9.44154 | N 44.19654<br>E 9.58915     | N 44.19654<br>E 9.58915 | N 44.19811<br>E 9.58860 | N 44.53202<br>E 88.181 | N 44.53202<br>E 88.181         | N 44.58542<br>E 6.92014 | N 44.58542<br>E 6.92014 | N 44.40907<br>E 8.54205 | N 44.40907<br>E 8.54205 | N 44.55364<br>E 8.44636 | N 44.55364<br>E 8.44636 |        |                                |  |  |  |
| SiO <sub>2</sub>                     | 37.80                   | 0.02                    | 25.40                   | 11.46                   | 25.28                   | 16.63                   | 34.53                       | 19.54                   | 14.04                   |                        |                                | 6.96                    | 21.22                   | 6.13                    | 13.05                   | 18.66                   | 21.26                   | 12.59  |                                |  |  |  |
| Al <sub>2</sub> O <sub>3</sub>       | 1.69                    | 0.01                    | 1.09                    | 0.65                    | 1.57                    | 1.27                    | 1.10                        | 1.71                    | 0.54                    |                        |                                | 0.16                    | 1.14                    | 0.24                    | 0.25                    | 1.34                    | 1.03                    | 0.52   |                                |  |  |  |
| Fe <sub>2</sub> O <sub>3</sub> (Tot) | 8.37                    | 0.03                    | 5.14                    | 3.43                    | 5.92                    | 5.56                    | 6.53                        | 5.26                    | 3.17                    |                        |                                | 1.20                    | 3.28                    | 1.59                    | 2.92                    | 8.47                    | 4.97                    | 2.67   |                                |  |  |  |
| MnO                                  | 0.123                   | 0.080                   | 0.071                   | 0.080                   | 0.098                   | 0.093                   | 0.124                       | 0.092                   | 0.100                   |                        |                                | 0.042                   | 0.124                   | 0.114                   | 0.121                   | 0.413                   | 0.063                   | 0.056  |                                |  |  |  |
| MgO                                  | 37.14                   | 0.17                    | 15.75                   | 10.29                   | 21.34                   | 10.45                   | 31.57                       | 14.38                   | 12.90                   |                        |                                | 6.16                    | 14.71                   | 4.81                    | 14.41                   | 25.82                   | 19.58                   | 11.51  |                                |  |  |  |
| CaO                                  | 0.12                    | 57.73                   | 26.35                   | 39.95                   | 20.77                   | 34.91                   | 8.42                        | 30.52                   | 35.96                   |                        |                                | 46.51                   | 31.46                   | 47.96                   | 34.57                   | 16.28                   | 26.28                   | 38.67  |                                |  |  |  |
| Na <sub>2</sub> O                    | < 0.01                  | 0.02                    | 0.07                    | 0.01                    | < 0.01                  | 0.03                    | < 0.01                      | 0.03                    | < 0.01                  |                        |                                | < 0.01                  | 0.01                    | < 0.01                  | < 0.01                  | 0.04                    | < 0.01                  | < 0.01 |                                |  |  |  |
| K <sub>2</sub> O                     | < 0.01                  | < 0.01                  | < 0.01                  | < 0.01                  | 0.02                    | < 0.01                  | < 0.01                      | < 0.01                  | < 0.01                  |                        |                                | < 0.01                  | < 0.01                  | < 0.01                  | < 0.01                  | 0.01                    | < 0.01                  | < 0.01 |                                |  |  |  |
| TiO <sub>2</sub>                     | 0.045                   | 0.020                   | 0.026                   | 0.013                   | 0.030                   | 0.013                   | 0.023                       | 0.051                   | 0.007                   |                        |                                | 0.003                   | 0.036                   | 0.013                   | 0.008                   | 0.043                   | 0.019                   | 0.007  |                                |  |  |  |
| P <sub>2</sub> O <sub>5</sub>        | < 0.01                  | < 0.01                  | 0.02                    | 0.03                    | 0.02                    | 0.02                    | 0.03                        | 0.03                    | 0.03                    |                        |                                | 0.01                    | < 0.01                  | < 0.01                  | 0.06                    | 0.04                    | 0.03                    | 0.03   |                                |  |  |  |
| H <sub>2</sub> O*                    | 13.32                   | < 0.01                  | 5.58                    | 4.43                    | 8.22                    | 2.30                    | 10.46                       | 6.10                    | n.d.                    |                        |                                | 3.51                    | 4.71                    | 0.92                    | 7.70                    | 15.44                   | 6.36                    | 2.75   |                                |  |  |  |
| CO <sub>2</sub>                      | < 0.01                  | 41.88                   | 19.50                   | 30.00                   | 16.50                   | 28.00                   | 7.00                        | 22.30                   | n.d.                    |                        |                                | 35.20                   | 23.20                   | 38.40                   | 26.40                   | 14.10                   | 19.90                   | 31.10  |                                |  |  |  |
| LOI                                  | 13.32                   | 41.88                   | 25.08                   | 34.43                   | 24.72                   | 30.30                   | 17.46                       | 28.40                   | 32.97                   |                        |                                | 38.71                   | 27.91                   | 39.32                   | 34.10                   | 29.54                   | 26.26                   | 33.85  |                                |  |  |  |
| Total                                | 98.61                   | 58.12                   | 99.01                   | 100.30                  | 99.79                   | 99.28                   | 99.79                       | 100.00                  | 99.73                   |                        |                                | 99.75                   | 99.90                   | 100.20                  | 99.49                   | 100.70                  | 99.50                   | 99.89  |                                |  |  |  |
| Li                                   | 1.24                    | 9.94                    | 5.47                    | 4.19                    | 15.30                   | 2.46                    | 14.29                       | 10.57                   | 5.79                    | 3.68                   | 3.46                           | 12.13                   | 1.85                    | 2.90                    | 3.04                    | 0.16                    | 5.57                    |        |                                |  |  |  |
| B*                                   | 24.6                    | 22.6                    | n.d.                    | 12.2                    | 82.1                    | 5.4                     | 13.4                        | 42.9                    | n.d.                    | 17.4                   | 16.2                           | 21.1                    | 10.9                    | 8.6                     | 12.8                    | 18.9                    | 10.9                    |        |                                |  |  |  |
| Sc                                   | 14.9                    | 2.22                    | 7.13                    | 4.41                    | 8.42                    | 3.89                    | 9.39                        | 7.38                    | 3.26                    | 0.97                   | 1.06                           | 5.81                    | 1.43                    | 3.30                    | 5.36                    | 7.22                    | 2.44                    |        |                                |  |  |  |
| V                                    | 57                      | 18                      | 32                      | 24                      | 48                      | 30                      | 41                          | 22                      | 16                      | 5                      | 5                              | 29                      | 6                       | 12                      | 30                      | 30                      | 12                      |        |                                |  |  |  |
| Co                                   | 109.0                   | 1.7                     | 55.0                    | 23.0                    | 48.0                    | 52.0                    | 101.0                       | 53.0                    | 34.0                    | 20.0                   | 19.0                           | 48.0                    | 8.9                     | 27.7                    | 53.3                    | 49.0                    | 36.9                    |        |                                |  |  |  |
| Ni                                   | 2200                    | 13                      | 1179                    | 600                     | 1015                    | 1150                    | 2000                        | 880                     | 870                     | 350                    | 350                            | 790                     | 160                     | 430                     | 500                     | 950                     | 750                     |        |                                |  |  |  |
| Cu                                   | 44.0                    | 26.0                    | 39.0                    | 39.0                    | 45.0                    | 12.2                    | 17.0                        | 25.0                    | 23.0                    | 100.0                  | 13.6                           | 22.9                    | 15.5                    | 21.7                    | 22.0                    | 90.4                    | 17.8                    |        |                                |  |  |  |
| Zn                                   | 62.00                   | 33.00                   | 28.00                   | 33.00                   | 35.00                   | 32.17                   | 31.70                       | 40.00                   | 23.00                   | 49.00                  | 13.69                          | 23.94                   | 14.77                   | 18.46                   | 43.59                   | 40.53                   | 17.08                   |        |                                |  |  |  |
| Ga                                   | 1.443                   | 2.795                   | 0.974                   | 0.604                   | 1.204                   | 0.862                   | 1.119                       | 1.219                   | 0.428                   | 0.096                  | 0.106                          | 0.841                   | 0.215                   | 0.389                   | 2.759                   | 0.974                   | 0.296                   |        |                                |  |  |  |
| As                                   | 0.038                   | 1.202                   | 0.551                   | 0.212                   | 0.500                   | 1.072                   | 2.504                       | 1.054                   | 0.515                   | 0.109                  | 0.087                          | 0.803                   | 0.155                   | 4.071                   | 9.737                   | 0.339                   | 0.162                   |        |                                |  |  |  |
| Rb                                   | b.d.l.                  | 17.5                    | 0.412                   | 0.088                   | 0.958                   | 0.245                   | 0.458                       | 0.111                   | 0.158                   | 0.101                  | 0.076                          | 0.217                   | 0.244                   | 0.189                   | 0.274                   | 0.198                   | 0.067                   |        |                                |  |  |  |
| Sr                                   | 2.23                    | 1365                    | 86.1                    | 103                     | 97.1                    | 120                     | 32.1                        | 74.3                    | 68.9                    | 137                    | 155                            | 236                     | 525                     | 633                     | 212                     | 184                     | 117                     |        |                                |  |  |  |
| Y                                    | 1.370                   | 7.170                   | 0.665                   | 0.919                   | 1.213                   | 0.630                   | 0.571                       | 2.297                   | 0.438                   | 0.155                  | 0.149                          | 1.166                   | 1.550                   | 2.155                   | 32.438                  | 2.716                   | 0.291                   |        |                                |  |  |  |
| Zr                                   | 0.212                   | 14.640                  | 0.120                   | 0.057                   | 0.134                   | 0.477                   | 0.206                       | 2.068                   | 0.089                   | 0.171                  | 0.124                          | 1.215                   | 0.646                   | 0.194                   | 0.492                   | 0.407                   | 0.370                   |        |                                |  |  |  |
| Nb                                   | b.d.l.                  | 2.410                   | 0.008                   | b.d.l.                  | b.d.l.                  | 0.056                   | 0.012                       | 0.036                   | 0.001                   | 0.020                  | 0.018                          | 0.007                   | 0.026                   | 0.032                   | 0.621                   | 0.047                   | 0.009                   |        |                                |  |  |  |
| Cd                                   | 0.034                   | 0.447                   | 0.101                   | 0.245                   | 0.099                   | 0.227                   | 0.090                       | 0.180                   | 0.160                   | 0.072                  | 0.072                          | 0.184                   | 0.043                   | 0.168                   | 0.139                   | 0.073                   | 0.060                   |        |                                |  |  |  |
| Sb                                   | 0.004                   | 0.148                   | 0.092                   | 0.005                   | 0.019                   | 0.068                   | 0.034                       | 0.035                   | 0.040                   | 0.008                  | b.d.l.                         | 0.158                   | 0.018                   | 0.254                   | 0.269                   | 0.056                   | b.d.l.                  |        |                                |  |  |  |
| Cs                                   | b.d.l.                  | 1.14                    | 0.505                   | 0.022                   | 0.233                   | 0.107                   | 0.366                       | 0.515                   | 0.136                   | 0.526                  | 0.402                          | 0.763                   | 0.491                   | 0.128                   | 0.167                   | 0.055                   | 0.585                   |        |                                |  |  |  |
| Ba                                   | 0.08                    | 44.04                   | 4.12                    | 0.91                    | 1.65                    | 4.59                    | 1.60                        | 2.82                    | 3.04                    | 2.59                   | 2.50                           | 3.23                    | 2.57                    | 3.16                    | 3.27                    | 2.33                    | 3.06                    |        |                                |  |  |  |
| La                                   | 0.004                   | 7.477                   | 0.030                   | 0.092                   | 0.040                   | 0.213                   | 0.046                       | 0.445                   | 0.133                   | 0.061                  | 0.069                          | 0.150                   | 0.106                   | 0.815                   | 1.609                   | 0.609                   | 0.059                   |        |                                |  |  |  |
| Ce                                   | 0.010                   | 9.962                   | 0.019                   | 0.028                   | 0.008                   | 0.228                   | 0.072                       | 0.334                   | 0.030                   | 0.110                  | 0.097                          | 0.295                   | 0.345                   | 2.470                   | 5.498                   | 0.223                   | 0.078                   |        |                                |  |  |  |
| Pr                                   | 0.004                   | 1.426                   | 0.005                   | 0.008                   | 0.004                   | 0.035                   | 0.009                       | 0.093                   | 0.014                   | 0.014                  | 0.013                          | 0.064                   | 0.070                   | 0.460                   | 1.013                   | 0.110                   | 0.016                   |        |                                |  |  |  |
| Nd                                   | 0.046                   | 5.494                   | 0.033                   | 0.047                   | 0.036                   | 0.141                   | 0.046                       | 0.484                   | 0.070                   | 0.056                  | 0.048                          | 0.392                   | 0.426                   | 2.255                   | 5.640                   | 0.524                   | 0.063                   |        |                                |  |  |  |
| Sm                                   | 0.046                   | 1.031                   | 0.021                   | 0.021                   | 0.033                   | 0.035                   | 0.019                       | 0.150                   | 0.015                   | 0.014                  | 0.011                          | 0.131                   | 0.172                   | 0.460                   | 2.153                   | 0.105                   | 0.018                   |        |                                |  |  |  |
| Eu                                   | 0.021                   | 0.218                   | 0.011                   | 0.024                   | 0.012                   | 0.014                   | 0.017                       | 0.069                   | 0.010                   | 0.009                  | 0.006                          | 0.079                   | 0.155                   | 0.113                   | 0.508                   | 0.038                   | 0.009                   |        |                                |  |  |  |
| Gd                                   | 0.109                   | 1.123                   | 0.050                   | 0.050                   | 0.081                   | 0.052                   | 0.041                       | 0.225                   | 0.031                   | 0.020                  | 0.017                          | 0.192                   | 0.263                   | 0.430                   | 3.299                   | 0.188                   | 0.028                   |        |                                |  |  |  |
| Tb                                   | 0.024                   | 0.150                   | 0.010                   | 0.010                   | 0.018                   | 0.008                   | 0.009                       | 0.038                   | 0.004                   | 0.003                  | 0.003                          | 0.030                   | 0.042                   | 0.053                   | 0.619                   | 0.030                   | 0.005                   |        |                                |  |  |  |
| Dy                                   | 0.191                   | 0.884                   | 0.090                   | 0.081                   | 0.147                   | 0.061                   | 0.075                       | 0.270                   | 0.032                   | 0.021                  | 0.019                          | 0.185                   | 0.258                   | 0.277                   | 4.311                   | 0.218                   | 0.034                   |        |                                |  |  |  |
| Ho                                   | 0.047                   | 0.183                   | 0.022                   | 0.020                   | 0.036                   | 0.015                   | 0.018                       | 0.062                   | 0.008                   | 0.004                  | 0.004                          | 0.038                   | 0.047                   | 0.053                   | 0.979                   | 0.056                   | 0.008                   |        |                                |  |  |  |
| Er                                   | 0.146                   | 0.516                   | 0.074                   | 0.068                   | 0.117                   | 0.048                   | 0.064                       | 0.191                   | 0.026                   | 0.014                  | 0.013                          | 0.104                   | 0.104                   | 0.153                   | 3.077                   | 0.184                   | 0.025                   |        |                                |  |  |  |
| Tm                                   | 0.024                   | 0.075                   | 0.012                   | 0.011                   | 0.019                   | 0.007                   | 0.011                       | 0.029                   | 0.004                   | 0.002                  | 0.002                          | 0.014                   | 0.010                   | 0.023                   | 0.476                   | 0.030                   | 0.004                   |        |                                |  |  |  |
| Yb                                   | 0.168                   | 0.469                   | 0.090                   | 0.074                   | 0.130                   | 0.053                   | 0.080                       | 0.194                   | 0.027                   | 0.015                  | 0.012                          | 0.092                   | 0.051                   | 0.144                   | 3.267                   | 0.202                   | 0.029                   |        |                                |  |  |  |
| Lu                                   | 0.028                   | 0.067                   | 0.014                   | 0.013                   | 0.022                   | 0.008                   | 0.014                       | 0.032                   | 0.005                   | 0.002                  | 0.002                          | 0.015                   | 0.006                   | 0.021                   | 0.513                   | 0.032                   | 0.005                   |        |                                |  |  |  |
| Hf                                   | 0.026                   | 0.371                   | 0.012                   | 0.007                   | 0.016                   | 0.012                   | 0.009                       | 0.062                   | 0.004                   | 0.004                  | 0.004                          | 0.040                   | 0.020                   | 0.013                   | 0.036                   | 0.016                   | 0.008                   |        |                                |  |  |  |
| Ta                                   | b.d.l.                  | 0.142                   | 0.001                   | b.d.l.                  | b.d.l.                  | 0.003                   | 0.001                       | 0.002                   | b.d.l.                  | 0.001                  | 0.001                          | 0.001                   | 0.001                   | 0.003                   | 0.042                   | 0.003                   | b.d.l.                  |        |                                |  |  |  |
| Pb                                   | 0.084                   | 3.162                   | 2.269                   | b.d.l.                  | 0.101                   | 0.069                   | 0.089                       | 0.267                   | b.d.l.                  | 0.067                  | 0.136                          | 0.153                   | b.d.l.                  | 2.602                   | 1.410                   | 1.457                   | b.d.l.                  |        |                                |  |  |  |
| Th                                   | 0.001                   | 1.567                   | 0.002                   | 0.002                   | 0.001                   | 0.035                   | 0.005                       | 0.005                   | 0.001                   | 0.007                  | 0.011                          | 0.001                   | 0.011                   | 0.009                   | 0.046                   | 0.029                   | 0.001                   |        |                                |  |  |  |
| U                                    | 0.001                   | 0.740                   | 0.242                   | 0.197                   | 0.077                   | 0.158                   | 0.400                       | 0.071                   | 0.101                   | 0.234                  | 0.221                          | 0.044                   | 0.059                   | 0.052                   | 0.154                   | 0.083                   | 0.368                   |        |                                |  |  |  |

\* H<sub>2</sub>O calculated as difference between LOI (loss of ignition) and CO<sub>2</sub> measured

n.d. - not determined

b.d.l. - below detection limits

\* - data performed at TIMS

Table 2

[Click here to download Table: Table\\_2-rev.xlsx](#)**Table 2.** In situ major (in wt.%) and trace element (in ppm) composition of serpentine minerals

|                                    | <b>MNT 15-1</b>                |           |                              |           |                     |           | <b>MNT</b>      |
|------------------------------------|--------------------------------|-----------|------------------------------|-----------|---------------------|-----------|-----------------|
|                                    | <i>after pyroxene (n = 11)</i> |           | <i>after olivine (n = 8)</i> |           | <i>mesh (n = 3)</i> |           | <i>clast (n</i> |
|                                    | <i>sd</i>                      | <i>sd</i> | <i>sd</i>                    | <i>sd</i> | <i>sd</i>           | <i>sd</i> |                 |
| <b>SiO<sub>2</sub></b>             | 38.07                          | 0.91      | 40.35                        | 0.71      | 35.31               | 0.92      | 49.35           |
| <b>TiO<sub>2</sub></b>             | 0.10                           | 0.05      | 0.03                         | 0.02      | 0.03                | 0.00      | 0.02            |
| <b>Al<sub>2</sub>O<sub>3</sub></b> | 3.21                           | 0.92      | 0.89                         | 0.32      | 0.59                | 0.05      | 1.61            |
| <b>FeO</b>                         | 4.80                           | 1.08      | 5.04                         | 0.59      | 12.81               | 2.16      | 6.36            |
| <b>MnO</b>                         | 0.14                           | 0.02      | 0.08                         | 0.01      | 0.09                | 0.00      | 0.12            |
| <b>MgO</b>                         | 41.62                          | 0.77      | 41.55                        | 0.73      | 39.14               | 1.23      | 26.29           |
| <b>CaO</b>                         | 0.06                           | 0.02      | 0.06                         | 0.01      | 0.04                | 0.00      | 4.14            |
| <b>Na<sub>2</sub>O</b>             | <0.01                          |           | <0.01                        |           | <0.01               |           | 0.11            |
| <b>K<sub>2</sub>O</b>              | <0.01                          |           | <0.01                        |           | <0.01               |           | 0.60            |
| <b>Total</b>                       | 88.00                          |           | 87.98                        |           | 88.00               |           | 88.12           |
|                                    | <i>n = 10</i>                  |           | <i>n = 8</i>                 |           | <i>n = 3</i>        |           | <i>n =</i>      |
| <b>Li</b>                          | 2.0                            | 2.4       | 0.26                         | 0.18      | 0.061               | 0.037     | 9.3             |
| <b>B</b>                           | 50                             | 32        | 36                           | 11        | 32                  | 13        | 43              |
| <b>Sc</b>                          | 32                             | 12        | 11                           | 5         | 7.0                 | 0.8       | 3.6             |
| <b>V</b>                           | 152                            | 79        | 27                           | 23        | 16                  | 1         | 33              |
| <b>Cr</b>                          | 5900                           | 2000      | 700                          | 600       | 130                 | 70        | 1200            |
| <b>Co</b>                          | 13                             | 8         | 15                           | 4         | 86                  | 38        | 132             |
| <b>Ni</b>                          | 380                            | 150       | 1600                         | 400       | 2200                | 600       | 3100            |
| <b>Zn</b>                          | 12                             | 5         | 9.2                          | 1.5       | 8.8                 | 1.0       | 34              |
| <b>As</b>                          | 0.008                          | 0.002     | <0.060                       |           | 0.007               | 0.002     | 0.46            |
| <b>Rb</b>                          | 0.011                          | 0.005     | <0.023                       |           | 0.008               | 0.001     | 0.97            |
| <b>Sr</b>                          | 2.5                            | 1.4       | 2.2                          | 0.5       | 1.8                 | 0.1       | 22              |
| <b>Y</b>                           | 2.0                            | 0.8       | 1.7                          | 1.0       | 1.4                 | 0.2       | 0.083           |
| <b>Zr</b>                          | 0.27                           | 0.02      | 0.25                         | 0.05      | 0.17                | 0.03      | 0.11            |
| <b>Nb</b>                          | 0.020                          | 0.007     | <0.017                       |           | <0.001              |           | 0.013           |
| <b>In</b>                          | 0.018                          | 0.004     | <0.014                       |           | 0.006               | 0.001     | 0.012           |
| <b>Sb</b>                          | <0.004                         |           | <0.072                       |           | <0.004              |           | 0.045           |
| <b>Cs</b>                          | 0.010                          | 0.011     | <0.019                       |           | 0.002               | 0.000     | 1.5             |
| <b>Ba</b>                          | 0.032                          | 0.024     | <0.126                       |           | 0.028               | 0.004     | 4.2             |
| <b>La</b>                          | 0.002                          | 0.001     | <0.013                       |           | 0.001               |           | 0.012           |
| <b>Ce</b>                          | 0.012                          | 0.004     | <0.020                       |           | 0.007               | 0.000     | 0.007           |
| <b>Pr</b>                          | 0.006                          | 0.003     | <0.011                       |           | 0.003               | 0.000     | 0.005           |
| <b>Nd</b>                          | 0.082                          | 0.028     | 0.12                         | 0.03      | 0.052               | 0.005     | 0.079           |
| <b>Sm</b>                          | 0.078                          | 0.038     | 0.21                         | 0.07      | 0.053               | 0.012     | 0.13            |
| <b>Eu</b>                          | 0.036                          | 0.018     | <0.028                       |           | 0.025               | 0.005     | 0.051           |
| <b>Gd</b>                          | 0.19                           | 0.07      | 0.14                         | 0.03      | 0.12                | 0.02      | 0.067           |
| <b>Tb</b>                          | 0.039                          | 0.018     | 0.030                        | 0.011     | 0.025               | 0.003     | 0.081           |
| <b>Dy</b>                          | 0.33                           | 0.13      | 0.29                         | 0.10      | 0.21                | 0.03      | 0.29            |

Table 3

[Click here to download Table: Table\\_3-rev.xlsx](#)**Table 3.** In situ major (in wt.%) and trace element (in ppm) composition of carbonate minerals from

|                                    | MNT 13-5              |       |                     |       | ISC 13-1              |       |                     |       |
|------------------------------------|-----------------------|-------|---------------------|-------|-----------------------|-------|---------------------|-------|
|                                    | <i>clast</i> (n = 14) |       | <i>vein</i> (n = 5) |       | <i>clast</i> (n = 28) |       | <i>vein</i> (n = 4) |       |
|                                    | <i>sd</i>             |       | <i>sd</i>           |       | <i>sd</i>             |       | <i>sd</i>           |       |
| <b>SiO<sub>2</sub></b>             | <0.01                 |       | 0.04                | 0.02  | 0.21                  | 0.10  | 0.02                | 0.01  |
| <b>TiO<sub>2</sub></b>             | <0.01                 |       | <0.01               |       | 0.02                  | 0.02  | 0.03                | 0.03  |
| <b>Al<sub>2</sub>O<sub>3</sub></b> | <0.01                 |       | <0.01               |       | 0.01                  | 0.01  | 0.01                | 0.01  |
| <b>FeO</b>                         | <0.01                 |       | <0.01               |       | 0.17                  | 0.10  | 0.02                | 0.01  |
| <b>MnO</b>                         | 0.06                  | 0.04  | <0.01               |       | 0.20                  | 0.12  | 0.10                | 0.09  |
| <b>MgO</b>                         | 0.02                  | 0.01  | 0.54                | 0.22  | 0.51                  | 0.35  | 0.31                | 0.18  |
| <b>CaO</b>                         | 54.87                 | 2.54  | 54.42               | 0.61  | 56.02                 | 3.58  | 56.41               | 1.74  |
| <b>Na<sub>2</sub>O</b>             | 0.01                  | 0.01  | <0.01               |       | 0.02                  | 0.01  | 0.02                | 0.02  |
| <b>K<sub>2</sub>O</b>              | <0.01                 |       | <0.01               |       | <0.01                 |       | <0.01               |       |
| <b>Total</b>                       | 54.96                 |       | 55.00               |       | 57.15                 |       | 56.91               |       |
|                                    | <i>n</i> = 14         |       | <i>n</i> = 5        |       | <i>n</i> = 2          |       | <i>n</i> = 7        |       |
| <b>Li</b>                          | 0.75                  | 0.47  | 0.39                | 0.02  | 0.094                 | 0.030 | 0.45                | 0.09  |
| <b>B</b>                           | 0.48                  |       | 0.55                |       | 0.14                  | 0.05  | 0.93                | 0.05  |
| <b>Sc</b>                          | 0.44                  | 0.33  | 0.30                | 0.21  | 0.069                 | 0.008 | 1.6                 | 0.1   |
| <b>V</b>                           | 0.44                  | 0.34  | 0.20                |       | 0.009                 | 0.005 | 0.82                | 0.40  |
| <b>Cr</b>                          | 6.6                   | 2.8   | 2.7                 | 1.9   | 1.9                   | 0.2   | 11                  | 6     |
| <b>Co</b>                          | 1.2                   | 2.1   | 0.27                | 0.27  | 0.21                  | 0.14  | 4.7                 | 1.8   |
| <b>Ni</b>                          | 36                    | 19    | 11                  | 6     | 4.4                   | 2.7   | 320                 | 150   |
| <b>Zn</b>                          | 0.95                  | 0.45  | 0.16                | 0.09  | 0.26                  | 0.13  | 7.8                 | 3.0   |
| <b>As</b>                          | <0.014                |       | 0.083               | 0.108 | 0.021                 | 0.010 | 0.092               |       |
| <b>Rb</b>                          | 0.023                 | 0.012 | <0.007              |       | 0.006                 | 0.002 | 0.024               | 0.008 |
| <b>Sr</b>                          | 89                    | 26    | 160                 | 80    | 120                   | 50    | 120                 | 10    |
| <b>Y</b>                           | 0.42                  | 0.62  | 0.25                | 0.22  | 0.009                 | 0.004 | 0.12                | 0.02  |
| <b>Zr</b>                          | <0.008                |       | 0.010               |       | <0.027                |       | <0.001              |       |
| <b>Nb</b>                          | <0.005                |       | <0.004              |       | <0.003                |       | <0.002              |       |
| <b>In</b>                          | <0.004                |       | <0.003              |       | <0.014                |       | <0.014              |       |
| <b>Sb</b>                          | <0.020                |       | <0.017              |       | <0.075                |       | <0.003              |       |
| <b>Cs</b>                          | <0.004                |       | <0.004              |       | 0.026                 | 0.011 | 0.026               | 0.011 |
| <b>Ba</b>                          | 0.36                  | 0.18  | 0.39                | 0.35  | 1.5                   | 1.1   | 1.2                 | 0.3   |
| <b>La</b>                          | 0.014                 |       | 0.012               |       | 0.004                 | 0.002 | 0.019               |       |
| <b>Ce</b>                          | 0.012                 | 0.007 | 0.007               |       | 0.015                 | 0.001 | 0.015               | 0.001 |
| <b>Pr</b>                          | 0.011                 |       | <0.003              |       | 0.001                 |       | <0.001              |       |
| <b>Nd</b>                          | 0.056                 |       | 0.025               |       | <0.039                |       | <0.014              |       |
| <b>Sm</b>                          | 0.076                 | 0.047 | 0.062               |       | 0.004                 |       | <0.018              |       |
| <b>Eu</b>                          | 0.026                 |       | 0.010               |       | 0.002                 |       | <0.003              |       |
| <b>Gd</b>                          | 0.15                  | 0.15  | 0.040               |       | <0.037                |       | <0.013              |       |
| <b>Tb</b>                          | 0.053                 |       | 0.008               |       | <0.006                |       | <0.001              |       |
| <b>Dy</b>                          | 0.13                  | 0.16  | 0.029               | 0.004 | <0.020                |       | <0.007              |       |
| <b>Ho</b>                          | 0.045                 | 0.048 | 0.013               | 0.003 | 0.001                 |       | <0.001              |       |
| <b>Er</b>                          | 0.16                  | 0.15  | 0.044               | 0.019 | <0.023                |       | <0.010              |       |
| <b>Tm</b>                          | 0.047                 |       | 0.007               | 0.001 | <0.006                |       | <0.001              |       |

Table 4

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Table 4. Whole-rock C-O and Sr isotopic composition for serpentinite, limestone and ophicarbonates rocks

|                                       |                                   |                           |                           |         |                                 |          | <i>Jurassic age</i>                            |                           | <i>High-pressure age</i>                      |                          |
|---------------------------------------|-----------------------------------|---------------------------|---------------------------|---------|---------------------------------|----------|--|---------------------------|---|--------------------------|
|                                       |                                   | $\delta^{18}\text{O}$ (‰) | $\delta^{13}\text{C}$ (‰) | Rb/Sr   | $^{87}\text{Sr}/^{86}\text{Sr}$ | err      | $^{87}\text{Sr}/^{86}\text{Sr}_{160\text{Ma}}$ | $\Delta$ (measured-160Ma) | $^{87}\text{Sr}/^{86}\text{Sr}_{33\text{Ma}}$ | $\Delta$ (measured-33Ma) |
| <i>Oc. serp.</i>                      | <b>MNT 15-1</b>                   | <i>n.d.</i>               | <i>n.d.</i>               | -       | 0.706929                        | 0.000038 | 0.706929                                       | 0.000000                  |   |                          |
| <i>Limestone</i>                      | <b>ZRL 15-7</b>                   | 26.1                      | 0.5                       | 0.0128  | 0.707932                        | 0.000009 | 0.707848                                       | 0.000084                  |   |                          |
|                                       | <b>MNT 13-5</b>                   | 10.9                      | 1.4                       | 0.0048  | 0.706355                        | 0.000007 | 0.706323                                       | 0.000032                  |   |                          |
| <i>OCI (tectonic ophic.)</i>          | <b>ZRL 13-1</b>                   | 15.7                      | 0.6                       | 0.0009  | 0.706162                        | 0.000009 | 0.706156                                       | 0.000006                  |   |                          |
|                                       | <b>ZRL 15-3</b>                   | 15.6                      | 1.1                       | 0.0099  | 0.706768                        | 0.000016 | 0.706703                                       | 0.000065                  |   |                          |
|                                       | <b>ISC 13-1</b>                   | 16.2                      | 0.4                       | 0.0020  | 0.706510                        | 0.000012 | 0.706497                                       | 0.000013                  |   |                          |
| <i>OClI (sedimentary ophic.)</i>      | <b>SG 15-4</b>                    | 17.4                      | 1.1                       | 0.0143  | 0.705756                        | 0.000009 | 0.705662                                       | 0.000094                  |   |                          |
|                                       | <b>SG 15-2</b>                    | 15.9                      | -1.7                      | 0.0015  | 0.706046                        | 0.000022 | 0.706036                                       | 0.000010                  |   |                          |
|                                       | <b>MNT 15-2</b>                   | 17.0                      | 1.6                       | 0.0023  | 0.706790                        | 0.000009 | 0.706775                                       | 0.000015                  |   |                          |
|                                       | <b>PL 13-1</b>                    | 15.1                      | 0.3                       | 0.0007  | 0.706624                        | 0.000008 | 0.706619                                       | 0.000005                  |   |                          |
| <i>Prograde ophicarbonates</i>        | <b>PL 13-2</b>                    | 15.3                      | 0.2                       | 0.0005  | 0.706833                        | 0.000009 | 0.706830                                       | 0.000003                  |   |                          |
|                                       | <b>CU 3</b>                       | 17.7                      | -0.3                      | 0.0009  | 0.706383                        | 0.000008 | 0.706377                                       | 0.000006                  |   |                          |
|                                       | <b>CU 13-1</b>                    | 17.1                      | -2.8                      | 0.0005  | 0.707530                        | 0.000009 | 0.707527                                       | 0.000003                  |   |                          |
|                                       | <b>SA 15-1<sub>calcite</sub></b>  | 12.5                      | 1.0                       | <0.0001 | 0.707969                        | 0.000014 | 0.707969                                       | 0.000000                  | 0.707969                                      | 0.000000                 |
| <i>Eclogite facies ophicarbonates</i> | <b>SA 15-1<sub>dolomite</sub></b> | 11.9                      | 0.8                       | 0.0001  | 0.708130                        | 0.000022 | 0.708130                                       | 0.000000                  | 0.708130                                      | 0.000000                 |
|                                       | <b>SA 15-8</b>                    | 11.5                      | 0.4                       | 0.0013  | 0.709670                        | 0.000009 | 0.709661                                       | 0.000009                  | 0.709668                                      | 0.000002                 |
|                                       | <b>LP13 EOC1</b>                  | 11.3                      | 0.7                       | 0.0011  | 0.708453                        | 0.000005 | 0.708446                                       | 0.000007                  | 0.708452                                      | 0.000001                 |
|                                       | <b>LP13 EOC3</b>                  | 15.0                      | 0.5                       | 0.0006  | 0.706791                        | 0.000010 | 0.706787                                       | 0.000004                  | 0.706790                                      | 0.000001                 |

*n.d.* - not determined

**Supplementary Table A1**

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**Supplementary Table A2**

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**Supplementary Table A3**

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**Appendix Figure A1**

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**Supplementary material S1**

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