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Seismic properties of gabbroic sections in oceanic core complexes: constraints from seafloor drilling

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Abstract

The physical properties of rock within Oceanic Core Complexes (OCC) provide information about the history of magmatism, deformation and alteration associated with detachment faulting and unroofing of their gabbroic sequences. New core and logging data from Atlantis Bank (AB) OCC are compared with prior deep sea drilling data from this site on the Southwest Indian Ridge, as well as from Atlantis Massif (~~ME~~AM) OCC on the Mid-Atlantic Ridge. The average seismic velocity at each site is typical for gabbroic rock. Downhole core and log variations are found to depend on both porosity and alteration. Porosity at millimeter to meter scale impacts P-wave velocities (6.0–6.4 km/s) in the upper several hundred meters at both OCC. Below ~ 600–800 m depth, in situ velocities are higher (6.6–6.9 km/s) for the remainder of the 1.5 km drilled to date, except in zones where alteration is significant. Variability

of 0.5–1 km/s over depth intervals of a few meters is persistent where mineral alteration exceeds 10–20%, whereas variability drops to $\pm$ 0.1 km/s for intervals affected by little to no alteration. Seismic anisotropy is negligible overall, with just a few intervals showing consistent fast azimuth and magnitude of a few percent. The depth extent of alteration, probably related to the zone deformed below the detachment when active, is $\sim$ 600–800 meters at both AB and AM.

Keywords

Oceanic core complex

Seismic properties of ocean crust

International Ocean Discovery Program

JOIDES Resolution Expeditions 304, 305, 340T, and 360

Electronic supplementary material

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Introduction

Deep sea drilling provides unique access to ocean crustal properties at a scale that allows geologic processes to be addressed. The intrusive portion of the oceanic crust is typically inaccessible at the seafloor so seismic imaging and velocity analysis are applied to infer geologic structure of the basement. Averaging that is inherent in seismic survey methods means that narrow intervals marking key aspects of the geologic evolution of the crust may not be recognized. Here we take advantage of drilling and borehole logging results at two oceanic core complexes (OCCs) to illustrate inherent physical properties and to assess what the downhole variability might indicate about how they formed and the extent to which seismic surveys can and cannot detect such evidence of interplay between magmatism,

faulting, and fluid circulation.

OCCs expose intrusive crustal sections at/near the seafloor through sustained slip along detachment fault(s), which can be active for periods on the order of 10^6 year within the axial zone of slow spreading centers (e.g., Dick et al. 1991; Cann et al. 1997; Tucholke et al. 1998; Blackman et al. 2002; Escartin et al. 2008; Escartín et al. 2017; Macleod et al. 2009; John and Cheadle 2010). Understanding the inherent physical properties of the unroofed domal ‘core’ of these complexes can elucidate the footwall evolution during/following fault activity and illuminate aspects of the magmatic accretion that built the section that is penetrated by deep sea drilling.

AQ1

Two OCCs have been targeted for deep penetration by the Ocean Drilling Program (ODP, Legs 118, 176, 179), the Integrated Ocean Drilling Program (IODP Phase 1, Expeditions 304, 305, and 340T) and the International Ocean Discovery Program (IODP Phase 2, Expeditions 360 and 362T). Logging and selected shipboard description/measurement data for these deep penetrations are the main focus of this paper. ODP Hole 735B (Dick et al. 2000) and IODP Hole U1473A (MacLeod et al. 2017) were drilled into the footwall exposed at Atlantis Bank (AB), on the flank of the Southwest Indian Ridge (SWIR); IODP Hole U1309D (Blackman et al. 2011) was drilled into the Central Dome of Atlantis Massif (AM), on the flank of the Mid-Atlantic Ridge (MAR). AB developed as new crust accreted 11 Ma and the OCC subsequently rafted off axis to its current position ~ 80 km from the spreading center (Dick et al. 1991). AM developed 1–2 Ma and the east slope of its Southern Ridge currently forms the western rift valley wall near the ridge-transform intersection (Cann et al. 1997; Blackman et al. 2002). Regional seafloor morphology shows that detachment faulting and core complex formation in these areas is an episodic process (Fig. 1), with more typical faulting/magmatism developing linear abyssal hills in the intervening periods. At AB, transtension across the transform also contributed to the high-relief

morphology (Dick et al. 1991). Half spreading rate in these sections of the SWIR and MAR average 8 km/Myr (Dick et al. 1991; Hosford et al. 2003) and 12 km/Myr (Zervas et al. 1995), respectively.

Fig. 1

Ridge-transform areas where the AB (left) and AM (right), oceanic core complex occur, at the same scale. Brackets on AM map mark the Southern Ridge to the south of the Central Dome). Circles show location of drill holes and Site numbers for which core/log data are discussed. Gray arrows indicate relative plate motion at spreading axes (opposing arrows) and transform faults (single side arrow). Index map also shows locations (+) of Kane core complex (Atlantic) and Hess Deep tectonic window (Pacific) where gabbroic sequences were also recovered

Overall, core recovery from the OCCs was very good (Dick et al. 2000; Blackman et al. 2011; MacLeod et al. 2017) and logging produced a considerable amount of useful data for assessing physical properties. Hole 735B at AB extends 1508 m below the seafloor and the upper 500 m was logged during ODP Leg 118 (Robinson and von Herzen 1989). A suite of gabbroic rocks was recovered, with olivine gabbro dominant, intervals of oxide gabbros, and minor troctolite and diabase (Dick et al. 2000). Hole U1473A, 2.2 km NNE of Hole 735B, penetrated 808 m through a similar gabbroic section at AB, and it was logged during IODP Expedition 360.

Both oxide- and olivine-gabbro are also prevalent in the 157 m deep Hole 1105A drilled at AB during ODP Leg 179, about a km NE of Hole 735B (Pettigrew et al. 1999). At AM, Hole U1309D extends 1415 m into the domal core of the complex and it was logged in stages during IODP Expeditions 304, 305, and 340T (Blackman et al. 2006; Expedition 340T Scientists 2012). The recovered core is somewhat more variable than at AB and consists of gabbro (56%), olivine gabbro (25%) and several percent each of olivine-rich troctolites, troctolites, oxide gabbros, and diabase. Details of the physical property logging and corresponding core sample results for Holes 735B and U1309D have been published previously (Iturrino et al. 2002; Collins et al. 2009; Blackman et al. 2014; Harding et al. 2016) and we summarize those findings below. New AB results from Hole U1473A are reported in more detail. Also, new determinations of U1473A and U1309D (at AM) core sample velocities at pressures up to 200 MPa are reported. Following the description of the logging and relevant core measurements, we compare and contrast the results and discuss possible implications for recognizing core complex evolution at AB and AM.

Observations

Hole U1473A, AB

Drilling of Hole U1473A in AB was challenging due to several fractured intervals where borehole stability and recovery were poor (MacLeod et al. 2017). Despite the challenges, core recovery was good overall, 59%, and logging of the borehole was accomplished over essentially the full ~ 800 m depth attained during Expedition 360 (Fig. 2). Physical properties of the recovered rock were obtained by shipboard measurements (room/Standard Temperature and Pressure, STP) on 8 cm³ cubes cut from representative petrologic units throughout the section (MacLeod et al. 2017); we refer to these sample porosity, density, and P-wave velocity (V_p) data in this paper. Sonic logging runs included a pass in cross dipole mode to assess seismic anisotropy of the section, in addition to velocity. Triple combo logging

runs included wallrock density, resistivity and borehole caliper measurements which are discussed here, as well.

Fig. 2

Downhole data for IODP Hole U1473A at AB, SWIR. Lithology determined by visual core description, color code at right. Logging runs show borehole width and condition (black: single caliper triple combo run; red/magenta: 2-caliper FMS run1; blue/cyan- FMS run3). Right two panels show 2-m running average of wallrock density and seismic velocity

AQ2

Overall, the wallrock density and V_p are typical for gabbros-
 $2850\text{--}2900\text{ kg/m}^3$ and $6\text{--}7\text{ km/s}$, respectively. Downhole velocity
 variability of $\pm 0.5\text{ km/s}$ is common over $5\text{--}10\text{ m}$ intervals in the upper $\sim$
 580 m of the section and this is probably real, based on the sufficiently

good borehole condition away from recognized fractured zones. Velocities are generally lower, ~ 6.4 km/s, at depths less than 580 mbsf, whereas V_p averages ~ 6.9 km/s below that depth. Mean V_p/V_s ratio for the full section is 1.8 ± 0.08 , with variability about the mean dropping from ± 0.1 above 580 mbsf to ± 0.03 at greater depths.

The presence of alteration veins correlates with the amplitude of density and velocity variability, as expected given the observed (MacLeod et al. 2017) clay, carbonate, or secondary amphibole fill compositions. The mean thickness of these veins averages 0.13 cm, with just two instances where a vein exceeds 2 cm in the recovered core. Veins up to 2 cm thick were not uncommon in the upper 100 m; mean thickness of 0.3–1 cm is not uncommon in the 100–300 mbsf interval. Below that, almost all veins in the recovered core were less than 0.2 cm in mean thickness. Most of the core from shallower than 580 mbsf was characterized as having an intensity of 1–5 veins per 10 cm. Below that depth, typical intensity drops to < 1 vein per 10 cm, except for an increase at 630–650 mbsf, below which the section is free of clay and carbonate veins (MacLeod et al. 2017). In this deepest section, logged velocities are more constant and the variation that does occur coincides with isolated vein(s) identified in the core and/or logged as spikes in gamma radiation (5–7 gAPI relative to background level of < 2 gAPI). Density in this deepest region shows some relation to core sample oxide content, although it is not simple. Oxide content was typically estimated by macroscopic core description, with occurrences of locally increased concentration guided by along-core magnetic susceptibility plots. Density values somewhat greater than 2900 kg/m^3 are typical from 580 to 685 mbsf, they drop below 2900 kg/m^3 for the olivine gabbro interval 685–710 mbsf, and then increase to almost 3000 kg/m^3 where oxide content is notable, to ~ 740 mbsf (Fig. 2). This downhole pattern is not mirrored in V_p , which would not be surprising if oxide dominates the density signal, since oxide gabbro velocity is lower than that of olivine gabbro (Iturrino et al. 2002).

The U1473A cross dipole sonic log data were processed to determine

anisotropy using the post-processing module in Schlumberger's Techlog software. Shear waves and their slowness are identified/determined by that algorithm, and rotation to obtain the polarization direction of the fast shear wave is solved. Not surprisingly, intervals that might have been most interesting in terms of anisotropy signature were those where fracturing degraded the borehole condition, so reliable results were not obtained. Outside of these zones, several intervals that are 10's m in length display relatively constant azimuth of fast shear wave propagation ($\sim N20^{\circ}W$, Figure S1) with low anisotropy magnitude $\sim 1\%$. Just a few meter-scale intervals indicate larger anisotropy, 3–4%, at subseafloor depths of 335, 410, and 545 mbsf. The two shallower intervals are associated/bounded by faults documented in the core so deformation-controlled porosity likely is responsible for the anisotropy. Between 541 and 548 mbsf, visual core description (MacLeod et al. 2017) indicates a fault, underlain by a zone of increased veining, both within olivine gabbro, which overlies oxide gabbro for which the upper contact was not recovered. So, deformation related orientation of porosity probably is responsible for the logged seismic anisotropy.

Core sample measurements at STP (MacLeod et al. 2017) document an expected general decrease in V_p as porosity increases in Hole U1473A but there is considerable scatter (Fig. 3a). Similarly, primary mineralogy is not a consistent factor in the measured core sample velocity, as indicated by the lack of clean trend of V_p with change in olivine mode (Fig. 3b), although olivine content does play a role. Gabbro samples (olivine mode < 5) have average V_p of 6.59 km/s whereas the average for olivine gabbros (mode 5–17 in the U1473A sample set) is 6.74 km/s. Alteration plays a more important role in reducing velocity for these samples, mean V_p values dropping on the order of 0.18 km/s for every 10% increase in overall alteration, as visually estimated (MacLeod et al. 2017).

Fig. 3

Hole U1473A core sample physical properties measured at STP (MacLeod et al. 2017). **a** Velocity shows broad decrease with increasing porosity, also

significant scatter. **b** Velocity shows very slight dependence on olivine mode in this 0–800 mbsf section

Hole U1309D, AM

Site U1309 in the footwall to the Central Dome detachment at AM comprises a gabbroic sequence that is probably ~ 4 km thick (Harding et al. 2016; Blackman and Collins 2010). Results of post-drilling seismic survey data analysis, employing **blue** downward continuation of streamer refraction data and waveform inversion, are consistent with the borehole data at the scale of imaging resolution (order 100–200 m, Harding et al. 2016). In the upper, drilled section, an overall increase in seismic velocity with depth is punctuated by local reductions, where alteration is locally high and/or deformation structures are more intense (Blackman et al. 2014). V_p increases from ~ 4 to 6 km/s from the seafloor to 350 mbsf, showing significant variability at meter(s) scale (Fig. 4). Average V_p is 6.0 km/s in the 350–800 mbsf interval where variation is still significant

(as much as ± 1 km/s). The lower part of the section, 800–1400 mbsf, has V_p of 6.6 km/s on average, with generally smaller meter-scale variability. The olivine-rich troctolite interval at 1080–1220 mbsf averages ~ 1 km/s slower than the surrounding section due to generally much greater alteration in that interval (Figs. 4 and 5), although local variation in the extent of serpentinization is significant (Fig. 5). The boundaries between petrologic units tend to show drops in V_p and, as expected, intervals where serpentinization is nearly complete (total alteration 80–100%) are associated with local dip in V_p .

Fig. 4

Downhole data for IODP Hole U1309D, at AM, MAR. Lithology determined by visual core description. Logging runs show borehole width/condition, and 2-m running average of wallrock density and seismic velocity. Exp 305 sonic log shown by purple curve and Exp 340T, 7 years later, is shown by red curve. Plot scales and lithology key are same as in Fig. 2

Fig. 5

Relation between primary and secondary lithology and P-wave velocity, shown for the deeper olivine-rich troctolite interval in Hole U1309D. Lithology shown at left is based on petrologic units described during Exp 304/305, as was the total alteration of core pieces (black symbols in right panel). Red curve shows the sonic log data with no downhole averaging.

Correspondence of depths between core and log may be uncertain by up to 1–2 m, still a common correlation between V_p variability and alteration is evident. Figure S2 shows the same data with depth scale expanded to aid core-log comparison

A fault at 746 mbsf has a damage zone extending at least 20 m below, and

possibly as much as that above (depending how much borehole degradation has impacted data quality) where V_p is 1–2 km/s lower than surrounding wallrock (Fig. 4). Low return and the presence of ultra-cataclasite (thin black interval in lithology column of Fig. 4) in the 17% that was recovered, along with analysis of images of borehole structure document this fault (Michibayashi et al. 2008). Diabase (gray at 760 mbsf in Fig. 4) intruded the section right at the base of the reduced V_p zone associated with this fault.

Cross dipole sonic log data for Expedition 340T covers the lower ~ half of the section and these data were processed the same way as the Hole U1473A data. For most of this lower section, the polarization direction of the fast shear wave is variable, a given value consistent only along intervals $< \sim 5$ m in length; the magnitude of anisotropy is low overall (generally $< 2\%$; Figure S1). Only in the 900–1080 mbsf interval, where a fairly uniform gabbroic sequence occurs, is there a constant fast azimuth ($N \sim 70^\circ W$) with anisotropy magnitude $\sim 1.5\%$. Two meter-scale intervals where anisotropy reaches 3–6% occur in the olivine-rich troctolite zone 1090–1200 mbsf.

In-situ velocity/density of OCC gabbroic rocks at lower crustal depth

Selected core samples from Hole U1309D were analyzed at pressures simulating conditions throughout the crust, to estimate in situ V_p for a multi-km thick gabbroic sequence. Results for 200 MPa ($\sim$ Moho depth for a standard ocean crustal section) were part of the seismic results reported by Collins et al. (2009). An error in the calibration used to compute those values was recognized recently, so we report corrected values here, along with showing how measured V_p on these samples varies as confining pressure is increased from 0 to 200 MPa. We analyze these results together with prior measurements made on gabbroic samples from ODP Holes 735B (AB), Hole 923A (70 m deep, Kane OCC) and [Hole 894G](#) (154 m deep, Hess Deep).

Hole U1309D rocks that were originally olivine-rich, either troctolite or olivine gabbro with more than 20% modal olivine, show a linear trend of increasing V_p (5.6–6.5 km/s) with increasing density (2650–2800 kg/m³, Fig. 6a). Most of the samples from this and the other sites scatter at higher value, in the ranges 6.5–7.5 km/s and 2800–3040 kg/m³, although the broad increase of V_p with density does continue the same basic trend as the originally olivine-rich samples. There is no clear pattern in V_p at 200 MPa with respect to gabbro sample depth. All but the U1309D highly altered olivine-rich samples, which occurred within the olivine-rich troctolite zone ~ 290–350 mbsf, vary throughout the 6.5–7.5 km/s range (Fig. 6b).

Fig. 6

Properties of core samples from gabbroic sections obtained by deep sea drilling, at room temperature. Density measured at room pressure and V_p measured at 200 MPa. Symbols/color are keyed to (I)-ODP Hole as indicated in legend on top panel. The lower density/velocity values correspond to a highly altered, olivine-rich troctolite interval in Hole U1309D. **a** V_p as a function of bulk density. **b** V_p versus depth in borehole

The increase in velocity with pressure was determined for each sample suite (Carlson and Miller 2004; Collins et al. 2009). Curves predicting the average increase in Vp with the natural log of estimated depth (Carlson and Miller 2004) were fit assuming an increase of 1 km per 20 MPa pressure increase. For Hole U1309D, gabbroic samples alone have average

velocity about 0.1 km/s higher than the full sample set tested (Fig. 7a). Predicted curves fit each sample set well and display ~ 0.5 km/s higher velocity in the upper 1–2 km of the crust that characterizes shallow gabbroic sequences, relative to the more typical upper crustal diabase (e.g. Hole 504B in crust formed at the East Pacific Rise), as shown in Fig. 7b.

Fig. 7 legend in Figure 7b has incorrect year in Cary & Chapman reference- it should be 1988 ...

Velocity of core samples as a function of depth, as measured in a series of lab runs at TAMU where confining pressure was increased incrementally and then values were measured again as pressure was incrementally decreased (Supplementary Material, Tables S1 and S2). **a** Corrected calibration for U1309D gabbro sample values with standard error shown, and curve fit solution for all rock types tested from that hole. **b** Curve fit solutions for sample suites from several gabbroic sections drilled by (I)ODP (735B, 894G, and 923A data from Carlson and Miller 2004). Seismic refraction data from various Atlantic studies are shown in gray curves, with crustal age and sources identified in legend. These selected expanding spread profile data have good velocity control and are located away from fracture zones

Differences between the velocity-depth curves for the different gabbroic sequences that have been drilled (Fig. 7b) indicate that alteration experienced by young intrusive oceanic crust would have a long-term velocity reduction effect as great as 0.25 km/s. Both the AM (U1309D) full dataset and the Hess Deep (894) dataset include more highly altered

samples than the other datasets and the curves that fit these data predict lower velocity. Results to date suggest that the thickness of the affected interval in young ocean lithosphere may be on the order of 800–1000 m, although drilling coverage is not currently sufficient to constrain the actual penetration of hydrothermal circulation and associated alteration. If there are fractures that extend to full crustal depths (several km) and there are intervals with olivine content 20% or greater, the velocity reduction due to associated seawater alteration would persist at depth as the plate ages.

Comparison of AB and AM

The core complex data from AM and AB display differences in physical properties, for which primary lithology appears to play some role but deformation and alteration history are controlling factors. The shipboard sample data (Fig. 8) indicate that seismic velocity is consistently lower at AM than at AB, despite bulk and grain density having significant overlap between the sites. Post expedition measurements of selected U1473A and U1309D samples using instrumentation aboard the *Chikyu* (Abe et al. 2018) confirm that the prior shipboard results do not simply reflect an unrecognized change in lab calibration on the *JOIDES Resolution*.

Fig. 8

Core sample porosity and Vp. **a** Comparison of prior (Iturrino et al. 2002; Blackman et al. 2006) and new (Macleod et al. 2017; Abe et al. 2018) properties at STP for AB and AM. **Red** curve should be red, not gray, in Figure 7a ... curve shows Vp:—Porosity prediction following formulation developed by (Carlson 2010; Carlson 2014); see Supplementary Material). **b** Core sample (black dot) and logged Vp (10-m average, red) in Hole U1309D. **(c–d)** Raw log (red, magenta) and sample Vp for 2 intervals at lab pressure of 0 (black dot), 10 (blue +), and 100 MPa (green o)

Porosity of the core samples is a major factor in the reduced V_p values at AM, compared to AB samples. Shipboard determinations of sample porosity are several percent higher in some of the AM section and this is one reason for the difference (Fig. 8a). Importantly, samples from deeper than ~ 800 mbsf in Hole U1309D show a steady decreasing V_p trend with depth, which is not observed in the borehole log (Fig. 8b). Hole 735B samples do not exhibit such a decrease in V_p for the lower part of the section (no log data are available for comparison at this site). We infer that Hole U1309D samples are currently under a combination of lithostatic and tectonic loads which, when released through drilling and recovery, results in microcracking that reduces V_p from the in situ value. Such effect has been noted in prior investigations of deep, hard rock core from the Pacific (Carlson, 2014). Current conditions at Site 735 in the 800–1500 mbsf interval must be such that sample unloading effects are minor. Larger

fractures affect in situ V_p at depths to several hundred meters below seafloor, where core sample values are almost always higher than log values (Fig. 8c, U1309D; Macleod et al. 2017, U1473A). These are not a factor at greater depths where lab V_p tracks logged value quite well (e.g. Fig. 8d; Macleod et al. 2017). Logged velocity is sensitive not only to effective properties that reflect distributed small cracks/porosity across the 2-m tool length, but also to features with size ranging up to many cm. Core samples, on the other hand, are chosen to avoid visible fractures and measurement of velocity on these cubic-2-cm pieces is sensitive to μm to sub-mm scale porosity.

In the upper ~ 800 m, differences in primary composition and/or the respective alteration histories probably also contribute to the reduced sample V_p at AM compared to AB. The latter is dominated by olivine gabbro (modal percent olivine $> 5\%$) whereas the upper kilometer and a half of AM comprises dominantly gabbro. Carlson et al. (2009) showed that the influence of olivine content on seismic velocities in gabbros depends on the $\text{Mg\#}$ as well as the degree of alteration to serpentine. In light of the comparatively high Mg numbers of the OCC samples (75–85), a greater olivine content would correspond to higher V_p in unaltered samples. While amounts are modest, AM contains several troctolite intervals and significant fractions of the olivine within these units are extensively altered, thereby lowering V_p overall despite (primary) olivine content being locally high.

The upper ~ 600 – 800 mbsf at both the AB and AM OCCs show greater variability in V_p than the underlying sections penetrated by seafloor drilling and this pattern probably reflects detachment-related processes (Fig. 9). The thickness of the deformation zone associated with detachment faulting at OCCs has been shown to be several hundred meters (Dick et al. 2000; John and Cheadle 2010; Blackman et al. 2014; Harding et al. 2016), with the intensity of impacts being greatest in the 200–400 m immediately beneath the fault (Schroeder and John 2004; Karson et al. 2006; Escartín et al. 2003). The association that we find between V_p variability, on the

meter-scale downhole, and the intensity of alteration veins indicates that seawater circulation is enabled by the detachment-induced deformation and footwall exposure. The fluid flow extends to depths of at least ~ 800 mbsf and persists long enough for reaction with the rock to leave a physically (and chemically) detectable record.

Fig. 9

Comparison of AB Holes 735B, U1473A, and AM Hole U1309D downhole data. Lithology shown as 20-m running average of recovered core. Alteration and deformation determined by visual core description on each expedition, with same intensity scale. All velocity depth curves are plotted together in center panel: dark green U1473A sonic log (10-m average); light green 735B sonic log (no ave, Iturrino et al. 2002); purple- U1309D sonic log (10-m ave.); gray- histogram of FullWave Inversion model within 1 km of U1309D (Harding et al. 2016); magenta- representative FWI curve for Southern Ridge of AM for comparison with Central Dome at Hole U1309D. Faulted intervals are marked by thick gray-green lines for U1473A and U1309D

Since AB is a wave-cut platform covered by carbonate deposits (Robinson and von Herzen 1989), it is unclear how much of the original section from just below the exposed detachment fault has been eroded away.

Relationships between the core recovered from Hole U1473A and Hole 735B and 1105A suggest that faulting within the bank may vertically offset these sequences by a up to a few hundred meters (MacLeod et al. 2017) so only rough depth correlation between the AB Sites is possible. Regardless, removal of a couple hundred meters of the detachment deformation zone could explain why cataclastic structures in the core from AB drop notably

at a shallower depth than at AM- somewhat < 600 mbsf versus ~ 800 mbsf, respectively. We cannot rule out the possibility that the detachment fault systems at AB and AM simply had different deformation distributions with depth. The occurrence of troctolites in the AM sequence and its proximity to a serpentinite-dominated domal core (the middle and western portion of the Southern Ridge, Henig et al. 2012; Harding et al. 2016; Fröh-Green et al. 2016), may indicate that magmatic conditions and possibly the interplay with faulting differed between the two OCCs (Escartín et al. 2003). Certainly, the extent of high temperature plastic deformation in the upper couple hundred meters of AB at Hole 735B is greater than that observed in the upper part of Hole U1309D (Ildefonse et al. 2007).

The similarity of depth to which Vp variability extends at AB and AM and the association of this variability with alteration veins suggests that effects of long-term aging on a ridge flank do not overwhelm signatures that develop as an OCC forms. Differences in the balance of magmatic/tectonic activity between the sites might be expected. ~~While~~ ~~Despite~~ AB ~~having~~ formed at an ultraslow spreading rate, Dick et al. (2000) conclude that melt supply in this segment center was higher than typical for most the SWIR. AM formed at the end of a slow-spreading segment and magmatism is interpreted to be episodic due to reduced melt supply far from the segment center (Blackman et al. 2011; Henig et al. 2012). Recent petrologic, geochemical, and structural analysis reported by Dick et al. (2019) support a fundamental difference in melt, crystallization, and accretion dynamics between AB and AM. However, while the associated differences in igneous stratigraphies might explain a modest offset in Vp between the sites, it cannot explain ~~a~~ the observed consistent depth extent of variability in Vp.

Borehole logs indicate that Vp of 6.9 km/s would likely characterize gabbroic bodies within the upper kms of OCC if faulting and alteration have not significantly affected them. For each of Holes 735B, U1473A, and U1309D, this value dominates the log at depths where cataclastic structures and alteration are generally quite low (< 1 vein per 10 cm, minor

fracturing). The uncertainty in the full waveform inversion also collapses for this part of the section at AM (Fig. 9, center panel), suggesting that advanced processing of multi-channel seismic data (Harding et al. 2016; Canales et al. 2008; Canales 2010; Xu et al. 2017) can provide some constraint on the extent of hydration within undrilled portions of OCC gabbro bodies.

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Electronic supplementary material

Below is the link to the electronic supplementary material.

Supplementary material 1 (DOCX 108 kb)

Supplementary material 2 (EPS 7759 kb)

Supplementary material 3 (EPS 3040 kb)

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